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Bachelor Thesis in Physics

submitted by

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GENERATION AND CONTROL OF X-RAY FREQUENCY COMBS THROUGH PERIODIC PHOTON MANIPULATION

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ABSTRACT

English

The possibility to generate and control x-ray frequency combs via periodic manipulation of the hyperfine magnetic field B in a nuclear forward scattering (NFS) setup is investigated theoretically. We consider NFS of x-ray pulses off a thin sample containing ^{57}Fe Mößbauer nuclei. These nuclei present a very narrow-band resonance at 14.4 keV. The magnetic field manipulation relies on periodic sequences of switching off and back on the B-field, which corresponds to coherent storage of the x-ray photon, and phase modulation by employing rotations by 180° of the B-field direction. Our numerical results show that x-ray combs with very narrow, equidistant peaks around the nuclear resonance energy can be generated and manipulated by tuning the parameters of the magnetic field switching sequence. Possible applications for metrology and quantum technology are anticipated.

Deutsch

Die Möglichkeit der Erzeugung und Kontrolle von Röntgen-Frequenzkämmen durch die Manipulation eines hyperfeinen magnetischen Feldes B in einem auf Kernresonanzstreuung in Vorwärtsrichtung (NFS) basierendem Aufbau wird theoretisch untersucht. Wir betrachten Röntgenpulse, die im Rahmen von NFS an einer dünnen Probe mit ^{57}Fe Mößbauer-Kernen streuen. Diese Kerne haben eine sehr schmale Resonanz bei 14.4 keV. Die magnetischen Schaltsequenzen basieren zum einen auf einem Aus- und wieder Anschalten des B-Feldes, wodurch eine kohärente Speicherung des Röntgenphotons induziert wird. Zum anderen verwenden die Sequenzen eine Rotation der Magnetfeldrichtung um 180°, welche zu einer Phasenmodulation des Photons führt. Unsere numerischen Resultate zeigen: Es ist möglich, sehr schmale, äquidistante Peaks nahe der Kernresonanzenergie zu generieren und zu manipulieren, indem man die Parameter der magnetischen Schaltsequenzen verändert. Mögliche Anwendungen werden in der Messtechnik und Quantentechnologie gesehen.

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CONTENTS

ACRONYMS

NFS   Nuclear Forward Scattering

QB   Quantum Beat

QM   Quantum Mechanics

qm   quantum mechanical

MBE   Maxwell-Bloch Equations

OBE   Optical Bloch Equations

SVE   slowly varying envelope

Inv   inverting sequence

NInv   non-inverting sequence

PDE   partial differential equation

Part I

PREPARATIONS

This part contains the general introduction and the derivation of the theoretical model employed. It initially provides a rough overview of this thesis' contents and subsequently explains the frame we work in.

INTRODUCTION

1.1 MOTIVATION

One of the key players for the development of quantum technologies is the photon as a flying qubit. It has the potential to become the fastest information carrier of tomorrow [1]. Present applications employ infrared, microwave and optical photons for photonic devices such as quantum photonic circuits [2] and quantum memories [3, 4]. However, from the angle of miniaturization, the large wavelengths of these photons and their corresponding diffraction limit pose a bottleneck to the predictions of Moore's Law [5]. Moving to shorter wavelengths, e.g. x-ray photons, would drastically reduce the limitation on the spatial dimensions. Apart from the smaller diffraction limit, x-rays convince with their deep penetration power, robustness and spatial resolution which potentially reaches down to atomic scales.

In addition, prospects for possible applications in quantum information and quantum optics [6] are fuelled by the recent commissioning of coherent x-ray sources [7, 8], new optical elements in this wavelength regime [9, 10] and nearly lossless detection efficiencies.

In order to employ the x-ray photons as qubits, it is inevitable to be able to control their core properties. As the x-ray wavelengths range on the energy scale of inner-shell atomic transitions [11] and low-lying nuclear transitions [12, 13], the resonant interaction might be exploited to exert control on the photons. First experiments in this topic have investigated non-linear effects of x-rays interacting with atoms [14, 15] as well as collective phenomena in the resonant scattering off nuclei [12, 16, 17]. The advantages of using nuclear transitions include that nuclei are clean, well-isolated systems that provide long coherence times and can be found in high-density solid-state targets.

One of the photon properties sought to control is their propagation speed. Even if just on a timescale of several hundred nanoseconds, the temporal qubit

storage is a key tool.

As was shown pioneeringly in Ref. [18] in 1996, storage of the photonic energy may be achieved in a Nuclear Forward Scattering (NFS) setup (see section 1.2) using nuclear resonances. The principal effect in NFS is the temporal creation of a delocalized nuclei excitation (Nuclear Exciton). At resonance, instead of exciting one single particular nucleus, the photon generates a collective excitation, that decays coherently in the forward direction [19].

Since 1996, theoretical research into this topic revealed further control opportunities through the manipulation of the external magnetic field. E.g. in Ref. [20], it was recently proposed that the rotation of the field can be used as a control knob to perform logical operations with the polarization of the flying qubits. Additionally, in Ref. [1], a fully coherent storage conserving not just energy but also phase and polarization of the photon was presented along with a 180° photon phase modulation. Here, the storage is achieved by switching off a previously present magnetic field embedding the nuclei. Switching the field back on at a later time releases the photon again. Furthermore, a shift of 180° in the scattered photon phase can be induced by abruptly rotating the field by 180°, so that the quantization axis is inverted. Moreover, it turns out that storage could be even more robust and more flexible when using special layered structures that contain a nanometer-thick iron film [21].

It is within the framework of the two mechanisms of storage and phase modulation where the findings of this thesis originate. We exploit the fact that these mechanisms can be combined and applied repetitively in a switching sequence. Instead of sticking to one single storage process, we theoretically investigate the effects of performing storage and modulation in a periodic pattern.

To do so, we choose a specific NFS geometry (see fig. 2), in which a short x-ray pulse of 1 ns duration shines onto a ^{57}Fe Mößbauer target.

Two main types of periodic magnetic switching sequences are tested with a field strength of $B_0 = 34.4\ T$:

i) The magnetic field alternates between 0 and B_0, while its direction is not altered. We refer to this as the "non-inverting" sequence in the following.

ii) The direction of the field is rotated by 180° so that it alternates between $-B_0$, 0 and B_0. We subsequently label this sequence "inverting".

Since we expect the periodicity to have an impact on the frequency composition of the forward detected radiation, we monitor the frequency as well as the time spectrum response.

Our numerical simulations show: comb structures of equidistant peaks emerge in the frequency domain of the resonantly forward scattered x-ray radiation. Depending on the specific parameters of the switching series, the combs change their shape. The increasing of the storage time narrows the combs and manifolds the frequency peaks, while the overall intensity declines. The combs of the non-inverting and the inverting series are shifted against each other by half their equidistant separation.

The generation of these frequency combs bears two opportunities. Firstly, the obvious utility is the application as a measurement instrument in the usual sense of frequency combs. In particular, the peak width in the order of tens of neV suggests actual opportunities. Secondly, since the distinctive form of the combs can be controlled via the parameters of the magnetic switching, a utilization to encode information is imaginable!

This work is structured as follows. In the remaining section of this chapter, we give a brief introduction to the physical base supporting this thesis, the NFS mechanism. We present the concept of Nuclear Exciton and describe the experimental setup theoretically considered. In the subsequent chapter 2, we lay the groundwork to theoretically investigate the dynamics of a light pulse propagating through a resonant medium. We start from a general Hamiltonian that can be split into an unperturbed and a perturbed part. We proceed by choosing a semi-classical approach to describe the radiation-nucleus interaction and by specifying a four-level quantum system. We close by putting forward the Maxwell-Bloch Equations (MBE) specific to our ansatz.

As the two photon manipulation techniques presented in Ref. [1] play a central role in our investigation methods, we dedicate chapter 3 to the reproduction of these results. We also provide a detailed description of the observed time spectra.

The likely most important chapter is chapter 4. There, we put forward the numerical results of investigating the scattered light dynamics under the influence of a periodically changing external magnetic field. We investigate the

generation and behaviour of above mentioned frequency combs and discuss their experimental practicability.

Finally, we draw conclusions from our results and present an outlook for the generation and use of the x-ray frequency combs.

1.2 NUCLEAR FORWARD SCATTERING

The foundation of our whole model is the phenomenon of Nuclear Forward Scattering (NFS). It is characterized by two main traits:

- the scattering happens in the same direction as the incident light (forward),

- collective effects (superradiance) occur.

The centrepiece of NFS is the Mößbauer effect.

The German physicist Rudolf Mößbauer originally set out to investigate the probability of reabsorption of γ-rays within solid-state samples. In 1958, he unexpectedly made an impactful discovery: the recoilless nuclear absorption of γ-ray photons.

Consider a γ-photon impinging on a target of absorbent particles, e.g. nuclei. The photon's energy has to cover the natural transition energy $\hbar\omega_t$ within a deviation of the natural linewidth Γ. Additionally, it has to provide the energy lost in recoil E_{rec} in order to drive an excitation: $E_\gamma \stackrel{!}{=} \hbar\omega_t + E_{rec}$.

If we have a homogeneous crystal target, all nuclei require the same energy to be excited. Hence, for a sufficiently small linewidth, a re-absorption within the medium can only occur if the photon or nucleus initially and after first re-emission compensates E_{rec} with energy additional to E_γ.

Multiple resonant scattering would therefore be possible in case the movement of the nuclei, i.e. phonon oscillations, provided such energy. Or if the recoil was negligible with respect to the transition width.

As Mößbauer revealed for specific materials, in re-absorption effects not involving phonons, for some scattering processes E_{rec} is sufficiently small due to a distributive nature of the excitation: the photon recoil collectively and equally spreads among all nuclei bound in the crystal. With the crystal mass
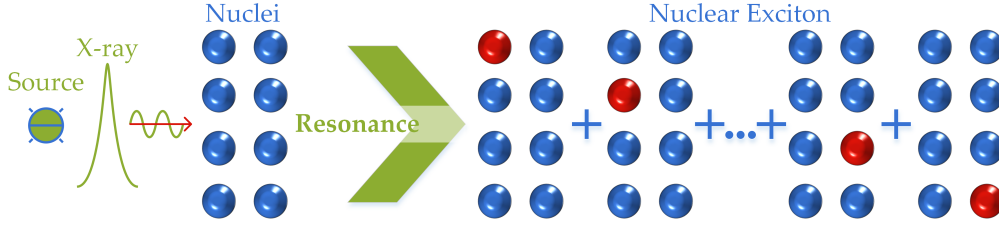
Figure 1: Schematic sketch of the Nuclear Exciton generation. A resonant light pulse (green) impinges on a crystal target (blue). The resulting excitation is delocalized and can therefore be described as a superposition of all possible single excitations (red).

being a macroscopic quantity, the recoil energy becomes negligible.

Starting from momentum conservation $|p_{rec}| \overset{!}{=} |p_\gamma|$, i.e.

$$\frac{E_\gamma}{c} = p_\gamma \overset{!}{=} p_{rec} = \sqrt{2ME_{rec}}, \tag{1.1}$$

we find for the recoil energy:

$$E_{rec} = \frac{E_\gamma^2}{2Mc^2}, \tag{1.2}$$

with the mass M of the collision partner.

The most popular Mößbauer nucleus is ^{57}Fe which has a mass of $m = 5.3 \cdot 10^7$ keV/c^2 and $E_\gamma = E_t = 14.4$ keV for the transition from the ground to the first excited state. We can find the natural linewidth $\Gamma = 2\Delta E$ from the uncertainty principle $\Delta E \Delta t \geq \hbar/2$ when setting the time imprecision to be the excited state's lifetime τ, which is $\tau = 141$ ns in the case of ^{57}Fe.

Herewith, the recoil/linewidth energy ratio is

$$\frac{E_{rec}}{\Gamma} = \frac{E_\gamma^2 \tau}{2M\hbar c^2} = \frac{1}{N}\frac{E_\gamma^2 \tau}{2m\hbar c^2} \approx \frac{4.3 \cdot 10^5}{N}, \tag{1.3}$$

with $M = N \cdot m$ as the mass of the crystal consisting of N single iron nuclei. We can see that for the Avogadro number $\frac{E_{rec}}{\Gamma} \approx 10^{-18}$ and for $N \to \infty$, the ratio converges to zero.

The recoilless absorption has a number of interesting and useful properties for the forward directed emission, i.e. photons scattered in their incidence direction:

- The lattice nuclei are kicked in a collective fashion. If additionally no

spin flip or other conversion occurs, there is no way of telling from the forward scattered radiation, which single nucleus was excited by the photon. The excitation is delocalized and it may be described as a superposition of all single nuclei. This is referred to as the "Nuclear Exciton" or "Timed Dicke State" in literature [22]. We may write its state as [23]

$$|\Psi_{NE}\rangle = \frac{1}{\sqrt{N}} \sum_{l=1}^{N} e^{i\mathbf{k_0 r_l}} |g_{\backslash l}\rangle |e_l\rangle$$

with $|g_{\backslash l}\rangle |e_l\rangle$ symbolizing a state in which the l-th nucleus at position r_l is excited (denoted by $|e_l\rangle$) while all other but l stay in the ground state ($|g_{\backslash l}\rangle$). The wave vector of the incident photon is labelled $\mathbf{k}_0$.

- Distinguishability in the time spectrum: Other scattering effects like Compton scattering or photoelectric effects are prompt, but the absorption-emission scattering of NFS is delayed by several nanoseconds [24]. The NFS response can therefore be distinguished simply by viewing the time spectrum from a certain delay point.

- Multiple scattering effects: For a large number of nuclei, the photon is resonantly and coherently scattered several times before leaving the target. The superposition of these excitations creates oscillations in the detected intensity (see chapter 3).

Setup

In our model (see chapter 2), we consider a ^{57}Fe solid-state sample. It is immersed in a magnetic field, determining the quantization axis to be y as shown in figure 2.
The target is penetrated by a short x-ray pulse (e.g. from a last generation synchroton radiation or x-ray free electron laser (XFEL) [25] source) propagating in z-direction. We tune the laser to drive the M1-transition of the Mößbauer nuclei between ground $|g\rangle$ and first excited state $|e\rangle$, which is at $\hbar\omega_0 = 14.4\ keV$. The nuclear spins are $I_g = 1/2$ and $I_e = 3/2$, allowing for a degeneracy via the magnetic quantum numbers $M_g = \pm 1/2$, $M_e = -3/2, \ldots, 3/2$. During periods when the magnetic field is present, the field lifts the degeneracy by separating the levels by their Zeeman splitting Δg, Δe for ground and excited state
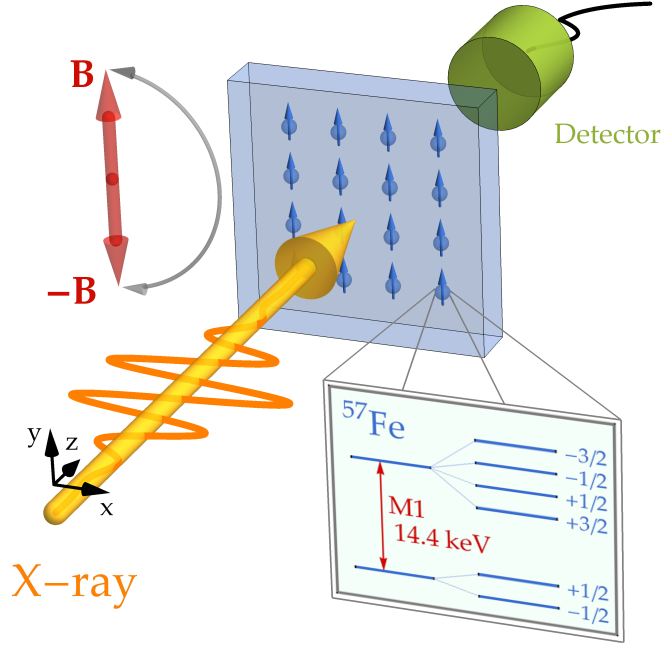
Figure 2: This figure shows the Nuclear Forward Scattering setup. The x-ray (orange) propagating along the z-direction penetrates the Mößbauer target. The probe experiences a magnetic field B, that can be switched off, on and rotated by π rad, so that the quantization axis is inverted. The inset shows the level scheme of the ground g and first excited e state of the ^{57}Fe sample including magnetic degeneration. The two states have the nuclear spins $I_g = 1/2$, $I_e = -1/2$, with magnetic quantum number degeneracies $M_g \in \{-1/2, 1/2\}$ and $M_e \in \{-3/2, -1/2, 1/2, 3/2\}$ as shown in the inset.
Many thanks to Jonas Gunst [27] for providing the image basis.

respectively. Since the pulse is wider than these splittings, it simultaneously addresses all possible six transitions which fulfil $\Delta M := M_e - M_g \in \{-1, 0, 1\}$. Following the steps in Ref. [26], we can find a way of describing the dependence of the scattered field amplitude on the setup geometry.

The coherently forward scattered radiation amplitude $E(z, t)$ can be determined from the iteratively solvable wave equation [28, 29]

$$\partial_z E(z, t) = -\sum_l K_l \, \mathbf{J}_l(t) \int_{-\infty}^{t} \mathbf{J}_l^\dagger(\tau) E(z, \tau) d\tau, \qquad (1.4)$$

with a constant K_l. The summation index l runs over all combinations of M_g, M_e for each nuclear site in the sample.

The nuclear transition current matrix elements $\mathbf{J}_l(t) = \langle I_g M_g \, | \, \mathbf{J}(t) \, | \, I_e M_e \rangle$ and $\mathbf{J}_l^\dagger$ represent excitation and de-excitation processes in the resonant scattering. The $\mathbf{J}$ operators are time-dependently proportional to the nuclear current den-

sity operator in momentum representation $\mathbf{j}(\mathbf{k})$, so that we can write

$$\mathbf{J}_l^\dagger(t) \sim \left\langle I_g M_g \left| \mathbf{j}^\dagger(\mathbf{k}) \right| I_e M_e \right\rangle = \mathbf{j}_l^\dagger(\mathbf{k}). \tag{1.5}$$

These density matrix elements are given for M1 transitions by [28, 30]

$$\mathbf{j}_l^\dagger(\mathbf{k}) = \sqrt{\frac{3(2I_e+1)c^5\Gamma_\gamma}{4\omega_0^3}} \sum_{q=0,\pm1} \begin{pmatrix} I_g & 1 & I_e \\ -M_g & q & M_e \end{pmatrix} (-1)^q \mathbf{k} \times \mathbf{n}_{-q}^*, \tag{1.6}$$

with the matrix denoting the Wigner 3-j symbol and the wave vector $\mathbf{k}$ of the incident light. We denote the speed of light by c, the M1-transition frequency by ω_0 and the radiative decay rate of the excited state by Γ_γ.

The $\mathbf{n}_q$ are spherical unit vectors that depend on the dictation of the quantization axis via the magnetic field. In Cartesian coordinates (with the canonical basis vectors $\mathbf{e}_i$), they are given by

$$\mathbf{n}_0 = \mathbf{e}_y,$$
$$\mathbf{n}_{+1} = -\frac{1}{\sqrt{2}}(\mathbf{e}_x + i\mathbf{e}_z), \tag{1.7}$$
$$\mathbf{n}_{-1} = \frac{1}{\sqrt{2}}(\mathbf{e}_x - i\mathbf{e}_z).$$

The vectorial orientation of the scattered field amplitude is given by its polarization $\mathbf{e}_0$, so that $\mathbf{E}_0(z,t) \sim \mathbf{e}_0$. Going back to Eq. (1.4), we see that the scattered amplitude for a single nuclear site and single $l = (M_g, M_e)$ is proportional to the scalar product

$$\mathbf{J}_l^\dagger(t)E(z,t) \sim \mathbf{j}_l^\dagger(\mathbf{k})\mathbf{e}_0$$

$$\sim \mathbf{e}_0 \left(\sum_{q=0,\pm1} \begin{pmatrix} I_g & 1 & I_e \\ -M_g & q & M_e \end{pmatrix} (-1)^q \mathbf{k} \times \mathbf{n}_{-q}^* \right)$$

$$\stackrel{*}{=} \mathbf{e}_0 \left(\begin{pmatrix} I_g & 1 & I_e \\ -M_g & (M_g - M_e) & M_e \end{pmatrix} (-1)^{(M_g - M_e)} \mathbf{k} \times \mathbf{n}_{(M_e - M_g)}^* \right), \tag{1.8}$$

where in $(*)$ we have used that the index q is determined by the 3-j symbol properties which demand $q = M_g - M_e$ in order for the symbol to not be zero [31]. Therefore $-q$ is equal to ΔM from above.

This last term in Eq. (1.8) vanishes for certain combinations of the polarization, wave vector and quantization axis orientations. This means that for specific geometries, some pairs $(M_g,\ M_e)$ do not contribute to the scattered field amplitude and the six level system's output is effectively reduced to a fewer level system's output.

To constrain our system in the inset in figure 2 to only four levels (see figure 3 on page 17), we set the x-ray pulse to be linearly polarized along the x-axis (so that $\mathbf{e}_0 = \mathbf{e}_x$). In combination with the displayed B-field geometry, the selection rule $\Delta M = 0$ kicks in [26]. Our four level system then consists of the $M_g = \pm 1/2$ and $M_e = \pm 1/2$ levels since only the two $\Delta M = 0$ transitions are driven by the laser.

2

THEORETICAL MODEL

In this chapter we derive the fundamental equations used in this thesis to describe the dynamic of a light pulse propagating through a target with resonant nuclei.

Our approach is based on a semi-classical description: while we treat the nuclei fully quantized, we approximate the radiation classically. This ansatz, despite eventually appearing controversial on first sight, has proven to agree very well with experimental results [18, 32–34].

Starting from the Schrödinger equation, we employ the density matrix formalism to describe our system via the Master Equation. The Master Equation provides us with the Optical Bloch Equations (OBE). Linking the interaction Hamiltonian to reduced transition probabilities $\mathcal{B}$, permits us to relate the OBE to the wave equation of the light beam. We thereby arrive at our final set of equations, the Maxwell-Bloch Equations (MBE). These coupled partial differential equations can then be solved numerically.

Since we later on investigate the case of linearly polarized light incident on ^{57}Fe Mößbauer nuclei (see fig. 2 on page 9), we can always find a B-field geometry where only the $\Delta M = 0$ transitions are driven. In this geometry the description of the light-nucleus interaction can be simplified to a four-level system as shown in figure 3.

In the following, we derive the MBE for such a system.

2.1 MASTER EQUATION

Pure and Mixed States

In Quantum Mechanics (QM), the evolution of a state $|\psi\rangle$ of a system is governed by the Schrödinger equation [35]

$$i\hbar\partial_t |\psi\rangle = \hat{H} |\psi\rangle,$$
(2.1)

where the states $|\psi\rangle$ are elements of a Hilbert space H.

If it is possible to express $|\psi\rangle$ as a linear combination of an eigenstate basis $B := \{|\phi_i\rangle\}_i$ of H, then $|\psi\rangle$ is called a pure state, i.e.

$$|\psi\rangle = \sum_i \alpha_i |\phi_i\rangle.$$
(2.2)

The expansion coefficients are given by the projection of the state onto the respective basis element, $\alpha_i = \langle \phi_i | \psi \rangle$. They represent the probability amplitudes of finding the eigenvalue of $|\phi_i\rangle$ when collapsing the wave function in a measurement [35].

This probability $p_{QM} = ||\alpha_i||^2$ does not represent a classical probability in the mathematical sense. Meaning that, before the measurement, the pure state $|\psi\rangle$ is not in just one eigenstate with a certain likeliness, but in a quantum mechanical superposition of all eigenstates at the same time.

Now, let us consider an ensemble of pure states $|\psi_i\rangle$, which may be from different Hilbert spaces. If we let the ensemble (independent from measurement) be in just one of those states with a certain likeliness p_i at the same time, we have a classical statistical ensemble with classical probabilities p_i.

The difference between p_i and p_{QM} is that the first one is based on uncertainty, i.e. a simple lack of information available to the observer, while the latter represents the genuine, counter-intuitive indetermination of qm objects.

Density Matrix Formalism

To describe such a statistical ensemble, also referred to as a mixed state, we may introduce the density matrix formalism [35]:

$$\hat{\rho} : H_1 \otimes H_2 \otimes ...H_m \rightarrow H_i, \ |\Phi\rangle \mapsto \hat{\rho} |\Phi\rangle, \quad (2.3)$$

where $\otimes$ denotes the tensor product, $\hat{\rho}$ is the density matrix itself and H_n are Hilbert spaces.

Making the definition

$$\hat{\rho} := \sum_i p_i |\psi_i\rangle \langle\psi_i|, \quad (2.4)$$

we can find the time evolution of our statistical ensemble from (2.1):

$$\partial_t \hat{\rho} = \partial_t \sum_i p_i |\psi_i\rangle \langle\psi_i| = \sum_i p_i \left(\partial_t |\psi_i\rangle \langle\psi_i| + |\psi_i\rangle \partial_t \langle\psi_i| \right)$$

$$= \frac{-i}{\hbar} \sum_i p_i \left(\hat{H} |\psi_i\rangle \langle\psi_i| + |\psi_i\rangle \hat{H} \langle\psi_i| \right)$$

$$= \frac{-i}{\hbar} \sum_i p_i \left[\hat{H}, |\psi_i\rangle \langle\psi_i| \right] = \frac{-i}{\hbar} \left[\hat{H}, \hat{\rho} \right]$$

$$\implies \partial_t \hat{\rho} = \frac{1}{i\hbar} \left[\hat{H}, \hat{\rho} \right]. \quad (2.5)$$

With (2.5), which is also known as the von Neumann equation, we have found a way of telling the state of our system $\hat{\rho}(t)$ at given times t.

The diagonal matrix elements $\langle\psi_i| \hat{\rho} |\psi_i\rangle =: \rho_{ii}$ are called populations for they describe the populations of the state $|\psi_i\rangle$, while the off-diagonals $\rho_{ij}, i \neq j$, are referred to as coherences since they give a measure of the coherence between $|\psi_i\rangle$ and $|\psi_j\rangle$.

So far, we have not specified the Hamiltonian of the system we are describing. In principal, Eq. (2.5) is qualified to fully describe the systems time evolution. However, its scope depends on the given Hamiltonian. It is possible, for example in the case of taking a semi-classical approach at the Hamiltonian, that fundamental processes are not included in the original von Neumann Equation.

To account for these neglected processes, we manually introduce the operator

$\hat{\rho}_s$:

$$\partial_t \hat{\rho} = \frac{1}{i\hbar}\left[\hat{H}, \hat{\rho}\right] + \hat{\rho}_s. \tag{2.6}$$

We will attend to the explicit form of the additional operator later, after specifying the Hamiltonian.

Equation (2.6) is the Master Equation and describes the evolution of our system, e.g. nuclei interacting with radiation. Its main constituent is the Hamiltonian $\hat{H}$, which is presented in section 2.2.

The equations of motion following from Eq. (2.6) are also referred to as the Optical Bloch equations. They only describe a single particle interaction, i.e. they do not take the response of an expanded, many-particle solid-state sample into consideration. This will be done in section 2.5 by including the Maxwell-response of the medium and obtaining the MBE.

2.2 THE HAMILTONIAN

Let us now consider an incident energy field driving transitions in a quantum system with discrete energies, like in our investigated setup shown in figure 2 on page 9.

Due to the selection rules given in chapter 1, the radiation-nucleus interaction is limited to the four-level scheme shown in figure 3. In the Schrödinger picture, the system's total Hamiltonian can be written as

$$\hat{H} = \hat{H}_0 + \hat{H}_1, \tag{2.7}$$

with an unperturbed part $\hat{H}_0$ and a perturbed one $\hat{H}_1$.

We assume the eigenvalues E_n of $\hat{H}_0$ to be known and hence write it as

$$\hat{H}_0 = \sum_n E_n \left|\psi_n\right\rangle \left\langle\psi_n\right| = \sum_n \hbar\omega_n \left|n\right\rangle \left\langle n\right|, \tag{2.8}$$

with angular frequencies ω_n and a simplified notation $\left|\psi_n\right\rangle = \left|n\right\rangle$ of the n-th eigenstate.
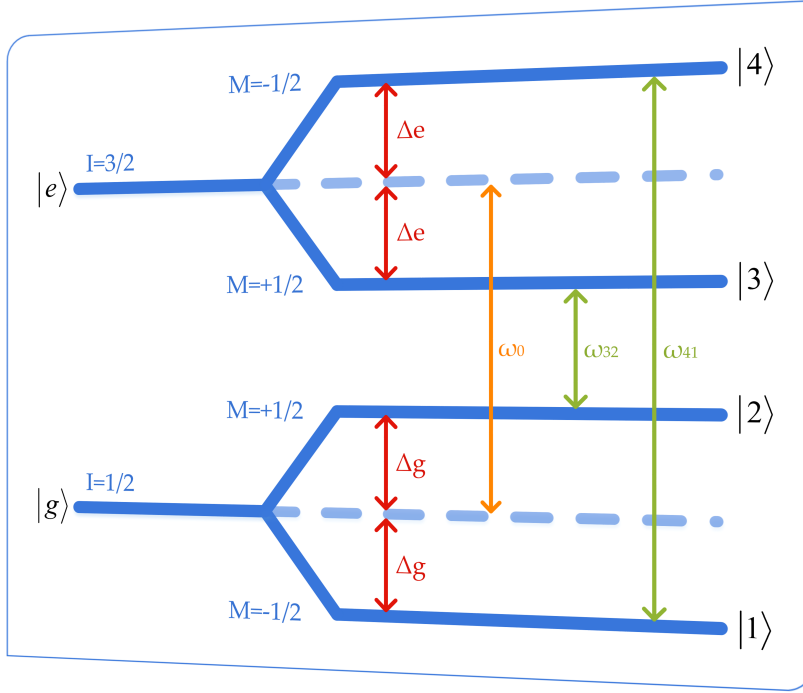
Figure 3: Schematic depiction of the four-level system in our ansatz. The ground $|g\rangle$ and first excited state $|e\rangle$ of ^{57}Fe originally have six (degenerate) energy levels. The respective Zeeman splittings are denoted by Δg and Δe. Under the constraint of linearly polarized light in a certain B-field geometry, only the two $\Delta M = 0$ transitions with $M = \pm\frac{1}{2}$ are driven. The transition energy from the excited to the ground state is $\omega_0 = 14.4 \, keV$.

The interaction part is given by the energy of the nuclear current $\hat{\mathbf{j}}$ and charge $\hat{\rho}_c$ in an external field, semi-classically described by a vector potential $\mathbf{A}$ and a scalar potential Φ [23]:

$$\hat{H}_1 = -\frac{1}{c}\hat{\mathbf{j}}(\mathbf{r}, t)\mathbf{A}(\mathbf{r}, t) + \hat{\rho}_c(\mathbf{r}, t)\Phi(\mathbf{r}, t). \tag{2.9}$$

To simplify this expression, we apply the Radiation Gauge for electro-magnetic fields (see Appendix A for the details). In this gauge, the scalar potential Φ is zero, so that

$$\hat{H}_1 = -\frac{1}{c}\hat{\mathbf{j}}(\mathbf{r}, t)\mathbf{A}(\mathbf{r}, t). \tag{2.10}$$

2.3 OPTICAL BLOCH EQUATIONS

Let us now derive the optical Bloch equations, which arise from the Master Equation (2.6).

Our system's dynamics are therein determined by

$$i\hbar\partial_t\rho_{ij} - \rho_{ij}^s = \left\langle i \left| \left[\hat{H}, \hat{\rho}\right] \right| j \right\rangle = \underbrace{\left\langle i \left| \left[\hat{H}_0, \hat{\rho}\right] \right| j \right\rangle}_{\text{①}} + \underbrace{\left\langle i \left| \left[\hat{H}_1, \hat{\rho}\right] \right| j \right\rangle}_{\text{②}}, \qquad (2.11)$$

with ρ_{ij}^s denoting the ij-th matrix element of the spontaneous decay matrix. The first right hand side term above equates to

$$\text{①} \stackrel{(2.8)}{=} \sum_k \underbrace{\left\langle i \left| \hat{H}_0 \right| k \right\rangle}_{=\delta_{ik}\omega_i} \left\langle k \left| \hat{\rho} \right| j \right\rangle - \left\langle i \left| \hat{\rho} \right| k \right\rangle \underbrace{\left\langle k \left| \hat{H}_0 \right| j \right\rangle}_{=\delta_{kj}\omega_j} = \rho_{ij}(\omega_i - \omega_j). \qquad (2.12)$$

For example from Fermi's Golden Rule we can see that the matrix element $\left\langle i \left| \hat{H}_1 \right| j \right\rangle$ of the interaction Hamiltonian is proportional to the transition rate from state $|j\rangle$ to $|i\rangle$.

Therefore we can conclude and define:

$$\left\langle i \left| \hat{H}_1 \right| j \right\rangle = \begin{cases} 0, & \text{i} \rightarrow \text{j is a forbidden transition} \\ -\frac{\hbar}{2}\Omega_{ij}, & \text{otherwise (relabelling).} \end{cases} \qquad (2.13)$$

The fraction $-\hbar/2$ was chosen to emphasize the matrix elements' connection to the Rabi frequency of a perturbed system.

In consequence, when considering the selection rule $\Delta M = 0$ according to our setup that allows one transition t per state, $|i\rangle \rightarrow |t(i)\rangle$, and the system defined as in figure 3, the second term reads:

$$\text{②} = -\frac{\hbar}{2}\left(\Omega_{it(i)}\rho_{t(i)j} - \rho_{it(j)}\Omega_{t(j)j}\right) \qquad (2.14)$$

with the map

i	$|1\rangle$	$|2\rangle$	$|3\rangle$	$|4\rangle$
t(i)	$|4\rangle$	$|3\rangle$	$|2\rangle$	$|1\rangle$

Seen that our final goal is to numerically solve the herefrom resulting equations, we conduct a phase transformation to get rid of high frequencies in the total Hamiltonian:

$$\rho_{ij} \rightarrow \rho_{ij}e^{i\omega_k t} =: \rho_{ij}' \qquad (2.15)$$

only for $(ij) \in \{(31), (41), (32), (42)\} =: Set$.

ω_k denotes the laser frequency.

This transformation is equivalent to making a change from the Schrödinger to an interaction picture frame [36].

If we define the unitary transform

$$U(t) := \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & e^{-i\omega_k t} & 0 \\ 0 & 0 & 0 & e^{-i\omega_k t} \end{pmatrix}, \tag{2.16}$$

we may obtain a transform into a rotating frame just as defined in Eq. (2.15). The density matrix then takes the form $\hat{\rho}' = U^\dagger(t)\,\hat{\rho}\,U(t)$, which equates to

$$\hat{\rho}' = \begin{pmatrix} \rho_{11} & \rho_{12} & e^{-i\omega_k t}\rho_{13} & e^{-i\omega_k t}\rho_{14} \\ \rho_{21} & \rho_{22} & e^{-i\omega_k t}\rho_{23} & e^{-i\omega_k t}\rho_{24} \\ e^{i\omega_k t}\rho_{31} & e^{i\omega_k t}\rho_{32} & \rho_{33} & \rho_{34} \\ e^{i\omega_k t}\rho_{41} & e^{i\omega_k t}\rho_{42} & \rho_{43} & \rho_{44} \end{pmatrix}. \tag{2.17}$$

Now, our ten equations read as follows:

I   for $(ij) \in Set$:

$$i\hbar\partial_t\rho'_{ij} - \rho^s_{ij} =$$
$$\rho'_{ij}(\omega_i - \omega_j - \omega_k) - \frac{\hbar}{2}\left(e^{i\omega_k t}\Omega_{it(i)}\rho_{t(i)j}\right.$$
$$\left. - e^{i\omega_k t}\Omega_{t(j)j}\rho_{it(j)}\right), \tag{2.18}$$

II   for $(ij) \notin Set$:

$$i\hbar\partial_t\rho_{ij} - \rho^s_{ij} =$$
$$\rho_{ij}(\omega_i - \omega_j) - \frac{\hbar}{2}\left(e^{sign(t(i)j)\omega_k t}\Omega_{it(i)}\rho_{t(i)j}\right.$$
$$\left. - e^{sign(it(j))\omega_k t}\Omega_{t(j)j}\rho_{it(j)}\right), \tag{2.19}$$

where

$$sign(t(i)j) := \begin{cases} +1, & \text{uneven} \\ -1, & \text{even permutation of (ij)} \in \text{Set.} \end{cases}$$

By studying the four-level system and the transition energies shown in figure 3, we can introduce the Zeeman splittings Δ_e, Δ_g as well as the laser detuning $\Delta := \omega_0 - \omega_k$ into our equation set. With ω_i denoting the frequency corresponding to the energy level of state $|i\rangle$, $i \in \{1, 2, 3, 4\}$, the explicit calculations are as follows:

- $\omega_2 - \omega_1 = 2\Delta g$,

- $\omega_3 - \omega_1 - \omega_k = \Delta g - \Delta e + \Delta$,

- $\omega_3 - \omega_2 - \omega_k = -(\Delta g + \Delta e) + \Delta$,

- $\omega_4 - \omega_1 - \omega_k = \Delta g + \Delta e + \Delta$,

- $\omega_4 - \omega_2 - \omega_k = \Delta e - \Delta g + \Delta$,

- $\omega_4 - \omega_3 = 2\Delta e$.

With this setup we have arrived at the final Optical Bloch Equations. From this point, we will omit the prime in Eq. (2.18), so that $\rho'_{ij} =: \rho_{ij}$.

2.4 THE MATRIX ELEMENTS

So far in our equations above, we are looking at ten unknown ρ_{ij} matrix elements as well as $\Omega_{ij} \propto \langle i | \hat{H}_1 | j \rangle$.

To better describe the latter, but more importantly to obtain an expression for the scattered field amplitude of the incident radiation, we will exploit the radial symmetry of the nuclei to expand the interaction matrix elements into spherical waves.

This will provide us with a characterization of the incident photons by their multipolarity L and its projection M onto the quantization axis, in equivalence to their wave vector $\mathbf{k}$ and polarization σ in a plane wave representation [27]. For this section, we mainly follow the arguments presented in Ref. [37] and

Ref. [38].

Initially, we expand the classical laser field in our perturbed Hamiltonian (section 2.2) in plane waves:

$$\mathbf{A} = \sum_k \mathbf{A}_k e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)} + c.c., \tag{2.20}$$

where we can drop the sum by restricting our scope to a single k mode.

From the Maxwell Equations in Radiation Gauge ($\Phi = 0$) we infer for the electric field $\mathbf{E}$:

$$\mathbf{E} = -\frac{1}{c}\partial_t \mathbf{A}. \tag{2.21}$$

Factorizing the fields into a slowly varying amplitude $\mathbf{A}_k(\mathbf{r},t)$ and a rapidly oscillating part, we may arrive at

$$\begin{aligned}
\mathbf{E}_k e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)} &= -\frac{1}{c}\left(-i\omega_k \mathbf{A}_k e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)} + e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)}(\partial_t \mathbf{A}_k)\right) \\
&\approx \frac{i\omega_k}{c}\mathbf{A}_k e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)}.
\end{aligned} \tag{2.22}$$

Accordingly, by writing the polarization vector $\mathbf{e}_{\mathbf{k}\sigma}^*$ explicitly, we have

$$\mathbf{A}(\mathbf{r},t) = \mathbf{A}_k(\mathbf{r},t)e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)} + c.c. \approx \frac{-ic}{\omega_k}E_k(\mathbf{r},t)\mathbf{e}_{\mathbf{k}\sigma}^* e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_k t)} + c.c. \tag{2.23}$$

From this we find for the matrix elements in Rotating Wave Approximation [23]

$$\begin{aligned}
\langle f\,|\,\hat{H}_1\,|\,i\rangle &= -\frac{1}{c}\left\langle f\,|\,\hat{\mathbf{j}}\cdot\mathbf{A}\,|\,i\right\rangle \\
&\approx \frac{i}{\omega_k}\int_V \mathbf{j}(\mathbf{r},t)\cdot E_k(\mathbf{r},t)\mathbf{e}_{\mathbf{k}\sigma}^* e^{i(\mathbf{k}\cdot\mathbf{r}-\Delta_k t)}\,d^3r + c.c. \\
&\overset{*}{\approx} \frac{i}{\omega_k}e^{-i\Delta t}\,E_k\int_V \mathbf{j}(\mathbf{r},t)\mathbf{e}_{\mathbf{k}\sigma}^* e^{i\mathbf{k}\cdot\mathbf{r}}\,d^3r + c.c.
\end{aligned} \tag{2.24}$$

where in (*) we have used that E_k is a slowly varying envelope and hence sufficiently constant over the vicinity of the integral, Δ is the transition-laser detuning.

Let us now proceed to expressing the integrand of Eq. (2.24) by the means of spherical waves.

Following the formalism described in Refs. [37], [38] and [39], we can for-

mulate the interaction matrix elements employing electric $\mathcal{E}$ and magnetic $\mathcal{M}$ multipole fields:

$$\mathbf{j}(\mathbf{r},t)\mathbf{e}_{\mathbf{k}\sigma}^{*}e^{i\mathbf{k}\cdot\mathbf{r}} = \sum_{L,M}\sqrt{2\pi(2L+1)}(-i)^{L}\,\mathcal{D}_{M,-\sigma}^{L}(\mathbf{k})\,\left(i\sigma\mathbf{A}_{LM}^{\mathcal{E}}(\mathbf{r}) + \mathbf{A}_{LM}^{\mathcal{M}}(\mathbf{r})\right),$$

(2.25)

with the Wigner rotation matrix $\mathcal{D}_{M,-\sigma}^{L}(\mathbf{k})$ [31] turning the quantization axis ($\hat{y}$) into the direction of $\mathbf{k}$.

Considering only the case of incident photons already parallel to the quantization axis, no rotation is needed, so that $\mathcal{D}_{M,-\sigma}^{L}(\mathbf{k}) = \delta_{M,-\sigma}$ now reduces to selecting the desired polarization, for which we can execute the sum over M.

The multipole fields $\mathbf{A}_{LM}^{\mathcal{E}/\mathcal{M}}(\mathbf{r})$ are given by

$$\mathbf{A}_{LM}^{\mathcal{E}}(\mathbf{r}) = \sqrt{\frac{L+1}{2L+1}}j_{L-1}(kr)\mathbf{Y}_{LL-1}^{M}(\theta,\phi) - \sqrt{\frac{L}{2L+1}}j_{L+1}(kr)\mathbf{Y}_{LL+1}^{M}(\theta,\phi)$$

$$= -\frac{i}{k}\nabla\times\left(j_{L}(kr)\mathbf{Y}_{LL}^{M}(\theta,\phi)\right)$$

and

$$\mathbf{A}_{LM}^{\mathcal{M}}(\mathbf{r}) = j_{L}(kr)\mathbf{Y}_{LL}^{M}(\theta,\phi).$$

(2.26)

$\mathbf{Y}_{LL}^{M}(\theta,\phi)$ are the vector spherical harmonics and $j_{L}(kr)$ the L-th spherical Bessel function.

Eq. (2.24) now reads

$$\langle f\,|\,\hat{H}_{1}\,|\,i\rangle \approx \frac{iE_{k}}{\omega_{k}}e^{-i\Delta t}\sum_{L}\sqrt{2\pi(2L+1)}(-i)^{L}$$

$$\times\int_{V}\mathbf{j}(\mathbf{r})\left(\frac{-i\sigma}{k}\nabla\times\left(j_{L}(kr)\mathbf{Y}_{LL}^{-\sigma}(\theta,\phi)\right) + j_{L}(kr)\mathbf{Y}_{LL}^{-\sigma}(\theta,\phi)\right)$$

(2.27)

Since the radiation wave length λ is much larger than the nuclear radius R_{n} the product $\frac{2\pi}{\lambda}R_{n} = kR_{n} \ll 1$ becomes sufficiently small to justify an expansion of $j_{L}(kr)$ in kr up to the first order. This is known as the long-wavelength approximation [38].

Involving the definitions of the electric $\mathbb{Q}_{LM}$ and the magnetic $\mathbb{M}_{LM}$ multipole

moment operators as [38]

$$\mathbb{Q}_{LM} = \int r^L Y^*_{LM}(\theta, \phi)\rho(\mathbf{r})d^3r,$$

$$\mathbb{M}_{LM} = \frac{1}{c(L+1)} \int \left[\mathbf{r} \times \mathbf{j}(\mathbf{r})\right] \nabla \left(r^L Y^*_{LM}(\theta, \phi)\right)d^3r, \tag{2.28}$$

our matrix elements take the form [27]

$$\langle f \,|\, \hat{H}_1 \,|\, i \rangle = -E_k e^{-i\Delta t}\sqrt{2\pi}\sum_L (-i)^{L+1}\sqrt{\frac{(2L+1)(L+1)}{L}}$$

$$\times \frac{\sigma k^{L-1}}{(2L+1)!!} \; \langle f \,|\, \mathbb{Q}_{L-\sigma} \,|\, i \rangle \tag{2.29}$$

in the case of electric transitions and

$$\langle f \,|\, \hat{H}_1 \,|\, i \rangle = -E_k e^{-i\Delta t}\sqrt{2\pi}\sum_L (-i)^{L}\sqrt{\frac{(2L+1)(L+1)}{L}}$$

$$\times \frac{k^{L-1}}{(2L+1)!!} \; \langle f \,|\, \mathbb{M}_{L-\sigma} \,|\, i \rangle \tag{2.30}$$

in the case of magnetic transitions.

Thereby, $n!!$ marks the double factorial, i.e. $n!! = n \cdot (n-2) \cdot (n-4) \cdot \ldots$.

So far, we reduced our interaction Hamiltonian to the calculation of the respective matrix elements of the electric and magnetic multipole moment operators. Exploiting that these operators are spherical tensors, we can apply the Wigner-Eckart theorem [31] to split off the dependency on the magnetic quantum number M, $-\sigma$:

$$\langle f \,|\, \mathbb{T}_{LM} \,|\, i \rangle = \frac{(-1)^{I_i - M_i}}{\sqrt{2L+1}}C(I_f I_i L; M_f - M_i M) \langle f \,\|\, \mathbb{T}_L \,\|\, i \rangle \tag{2.31}$$

for the M-th component of a spherical tensor $\mathbb{T}_{LM} = \mathbb{Q}_{LM}/\mathbb{M}_{LM}$ of rank L. The constants $C(j_1 j_2 J; M_1 M_2 M)$ denote the Clebsch-Gordan coefficients and I_i, M_i the spin of state $|i\rangle$ and its projection.

$\langle f \,\|\, \cdot \,\|\, i \rangle$ is commonly referred to as the reduced matrix element, despite not being an actual scalar product from the projection onto $\langle f|$ and $|i\rangle$ in the strict mathematical sense.

Now that our multipole operator matrix elements are independent of the angular momentum substructure, we can relate them to the principal transition

probabilities

$$\mathcal{B}(\mathcal{E}/\mathcal{M}L, I_i \to I_f) = \frac{1}{2I_i+1} |\langle f \,\|\, \mathbb{Q}_L/\mathbb{M}_L \,\|\, i\rangle|^2 \qquad (2.32)$$

and finally arrive at

$$\langle f \,|\, \hat{H}_1 \,|\, i\rangle = E_k e^{-i\Delta t}(-1)^{I_i-M_i}(-i)^{L+1}\sigma\sqrt{(2I_i+1)}C(I_f I_i L; M_f - M_i - \sigma)$$
$$\times \underbrace{\sqrt{\frac{2\pi(L+1)}{L}}\frac{k^{L-1}}{(2L+1)!!}\sqrt{\mathcal{B}(\mathcal{E}L, I_i \to I_f)}}_{=:\alpha(\mathcal{E}L,I_i\to I_f)}$$

$$(2.33)$$

in the case of electric transitions and constraining our system to only one multipolarity L.

We have chosen the definition of the multipole moment $\alpha(\mathcal{E}L, I_i \to I_f)$ as in Ref. [40].

With the magnetic transitions elements differing just by scaling or phase factors from the electric, we can summarize:

$$\langle f \,|\, \hat{H}_1 \,|\, i\rangle \sim E_k(z,t)\,(-1)^{I_i-M_i}C(I_f I_i L; M_f - M_i - \sigma)\,\alpha(\mathcal{E}/\mathcal{M}L, I_i \to I_f),$$

$$(2.34)$$

where we amended $E_k(\mathbf{r},t) = E_k(z,t)$ according to a unidirectional propagation along the z-axis.

2.5 THE WAVE EQUATION

As of now, we have neglected in our considerations the response the incident radiation generates in the expanded solid-state sample.

To include this in our model, we turn to the Maxwell equations in matter to find a wave equation for our radiation. As this equation imposes another condition for the sought solution, we aim to couple it to the other conditions (2.18) and (2.19). Doing so can be achieved by linking the wave equations source term to the density matrix and Rabi frequencies.

The (macroscopic) Maxwell equations read

$$i)\ div\mathbf{D} = 4\pi\rho \overset{\rho=0}{=} 0 \quad ii)\ div\mathbf{B} = 0$$

$$iii)\ rot\mathbf{E} = -\frac{1}{c}\partial_t\mathbf{B} \quad iv)\ rot\mathbf{H} = \frac{1}{c}\left(4\pi\mathbf{J} + \partial_t\mathbf{D}\right), \tag{2.35}$$

with the displacement and magnetizing fields $\mathbf{D} = \mathbf{E} + 4\pi\mathbf{P}$, $\mathbf{H} = \mathbf{B} - 4\pi\mathbf{M}$ and polarization $\mathbf{P}$ as well as magnetization $\mathbf{M}$ of the medium.

The vector $\mathbf{J}$ corresponds to the free macroscopic current density, i.e. the current induced by the light in the sample.

With $\mathbf{P}$, $\mathbf{M}$ set to zero, we obtain for the electric field [41]:

$$iii)\ \implies\ rotrot\mathbf{E} = -\frac{1}{c}\partial_t rot\mathbf{B}$$

$$\implies\ \left(-c^2\,\triangle + \partial_t^2\right)\mathbf{E} + 4\pi\mathbf{J} = 0, \tag{2.36}$$

or in one dimension

$$\left(\partial_z^2 - \frac{1}{c^2}\partial_t^2\right)\mathbf{E} = \frac{4\pi}{c^2}\mathbf{J}, \tag{2.37}$$

from which we can see, that the wave equation has $\mathbf{J}$ as source term.

Like above, we are assuming an electro-magnetic wave unidirectionally propagating in forward z-direction, which is now assumed to be linearly polarized along the x-axis:

$$\mathbf{E}(z,t) = E(z,t)e^{i(k\cdot z - \omega_k t)}\mathbf{e}_x, \tag{2.38}$$

with slowly varying envelope (SVE) E(z,t) and optical dispersion relation $k = \frac{\omega_k}{c}$.

Writing the current as $\mathbf{J}(z,t) = J(z,t)e^{i(k\cdot z - \omega_k t)}\mathbf{e}_x$, like proposed in Ref. [40], we can find for Eq. (2.37):

$$\left(\partial_z^2 + 2ik\partial_z - \frac{1}{c^2}\partial_t^2 + 2\frac{i\omega_k}{c^2}\partial_t\right)E(z,t) = \frac{4\pi}{c^2}\left(-i\omega_k + \partial_t\right)J(z,t)$$

$$\implies\ \left(\partial_z + \frac{1}{c}\partial_t\right)E(z,t) = -\frac{2\pi}{c}J(z,t). \tag{2.39}$$

Considering the slow propagation of the envelopes in time as well as in space, we note that for the respective highest order derivatives it is assumed that $|\partial_z^2 E| \ll |k\partial_z E|$, $|\partial_t^2 E| \ll |\omega_k\partial_t E|$ and $|\partial_t J| \ll |\omega_k J|$, which is why we applied the SVE approximation to retrieve Eq. (2.39).

According to Ref. [40], the current density for a single nuclear resonance may be obtained by summing over all nuclei involved in the coherent scattering and tracing over $\hat{\mathbf{j}}(\mathbf{k})e^{-ikz}\hat{\rho}$.

We exploit that, here, $\hat{\mathbf{j}}(\mathbf{k})$ is the nuclear deexcitation current operator in momentum space, meaning that the terms of $\left\langle j \,|\, \hat{\mathbf{j}} \,|\, i \right\rangle$ vanish or can be neglected except if $|i\rangle \rightarrow |j\rangle$ is a deexciting transition participating in the coherent scattering in the observed system [42].

Rewriting the source term accordingly yields:

$$
\begin{aligned}
J(z,t)\mathbf{e}_x = N \, tr(\hat{\mathbf{j}}e^{-ikz}\hat{\rho}) &= N \sum_{i=1}^{4} \left\langle i \left| \hat{\mathbf{j}}e^{-ikz}\hat{\rho} \right| i \right\rangle \\
&\stackrel{(*)}{=} N \sum_{i,h} \underbrace{\left\langle i \,|\, \hat{\mathbf{j}} \,|\, h \right\rangle}_{\text{selection rules}} \left\langle h \left| e^{-ikz}\hat{\rho} \right| i \right\rangle \\
&= N\left(\left\langle 1 \left| \hat{\mathbf{j}}e^{-ikz} \right| 4 \right\rangle \rho_{41} + \left\langle 2 \left| \hat{\mathbf{j}}e^{-ikz} \right| 3 \right\rangle \rho_{32} \right),
\end{aligned} \tag{2.40}
$$

where in (*) we have inserted the identity $\mathbb{1} = \sum_h |h\rangle \langle h|$ and N denotes the particle number density.

When taking into account the form of $\hat{H}_1$ from Eq. (2.24), but also in momentum space, $\hat{H}_1 = \frac{i}{\omega_k}\hat{\mathbf{j}}^*(\mathbf{k})E_k(z,t)\mathbf{e}_x e^{i(kz-\omega_k t)}$ [28], the wave equation fully connects to the optical Bloch equations by:

$$
\begin{aligned}
\left(\partial_z + \frac{1}{c}\partial_t \right) E_k(z,t) &= -\frac{2\pi}{c}J(z,t) \\
&= -\frac{2\pi}{c}N\left(\left\langle 1 \left| \hat{\mathbf{j}}e^{-ikz} \right| 4 \right\rangle \rho_{41} + \left\langle 2 \left| \hat{\mathbf{j}}e^{-ikz} \right| 3 \right\rangle \rho_{32} \right) \mathbf{e}_x \\
&= -\frac{2\pi\omega_k}{ic}\frac{N}{E_k(z,t)}\left(\left\langle 4 \,|\, \hat{H}_1 \,|\, 1 \right\rangle \rho_{41} \right. \\
&\quad + \left. \left\langle 3 \,|\, \hat{H}_1 \,|\, 2 \right\rangle \rho_{32} \right),
\end{aligned} \tag{2.41}
$$

where we phase-transformed $\hat{\rho}_{ij}$ as in Eq. (2.15) in section 2.3.

The wave equation now reads:

$$
\left(\partial_z + \frac{1}{c}\partial_t \right) \Omega(z,t) = \frac{i2\pi\omega_k}{\hbar c}N
$$
$$
\times \alpha^2(\mathcal{E}/\mathcal{ML}, I_g \rightarrow I_e)\left((-1)^{I_g - M_1}C_{41}\cdot\rho_{41} + (-1)^{I_g - M_2}C_{32}\cdot\rho_{32} \right) \tag{2.42}
$$

while employing the Clebsch-Gordan coefficients $C_{fi} := C(I_f I_i L, M_f - M_i M)$ as before and merging all matrix elements by defining $E_k(z,t)\alpha(\mathcal{E}/\mathcal{M}L, I_i \rightarrow I_f) = \dfrac{\Omega_{fi}}{\pm C_{fi}} =: \Omega(z,t)$.

Consider that α is independent of the magnetic degeneration, $\alpha(I_1 \rightarrow I_4) = \alpha(I_2 \rightarrow I_3) = \alpha(I_g \rightarrow I_e)$ with g, e denoting the ground and excited state in our setup (fig. 3).

Together with the OBE, we now accomplished the derivation of the MBE.

The values of the reduced transition probabilities $\mathcal{B}(\mathcal{E}/\mathcal{M}L, I_g \rightarrow I_e)$ are available from experimental determination.

For the purpose of convenience we now rewrite Eq. (2.42) following Ref. [40] so that we can replace $\mathcal{B}/\alpha$ by parameters we would like to manipulate in our setup.

$$\alpha^2(\mathcal{E}/\mathcal{M}L, I_g \rightarrow I_e) = \frac{\hbar\Gamma\sigma_r}{8\pi k}, \tag{2.43}$$

with the resonant cross section σ_r.

Exploiting the dispersion relation $\omega_k/k = c$, introducing the sample's optical thickness $\xi := N\sigma L/4$ and defining $\eta := \xi\Gamma/L$, we obtain the final wave equation in the simplified form

$$\left(\partial_z + \frac{1}{c}\partial_t\right)\Omega(z,t) = i\eta\left(C_{32} \cdot \rho_{32}(z,t) - C_{41} \cdot \rho_{41}(z,t)\right). \tag{2.44}$$

We also implemented $\left|I_g M_1\right\rangle = \left|1/2\,{-}1/2\right\rangle$, $\left|I_g M_2\right\rangle = \left|1/2\,1/2\right\rangle$ specific to our ^{57}Fe system to evaluate $(-1)^{I_g - M_g}$.

2.6 DECOHERENCE MATRIX

As mentioned in sec. 2.1, in the general case, the occuring decoherence processes that are not included in the Hamiltonian, may be accounted for by the introduction of a decoherence matrix $\hat{\rho}_s$. Let us now see, what this means for our specific system.

Phenomenologically, one can observe a spontaneous decay of the populations of the excited states and coherences [41]. Spontaneous emission depends upon quantum mechanical vacuum fluctuations and cannot be explained by classi-

cal physics.

Since we chose a classical description of the incident radiation, these processes are neglected by our Hamiltonian. We therefore set $\hat{\rho}_s$ to account for the spontaneous decay in our system.

As illustrated in figure 4, the populations of the excited states ρ_{44}, ρ_{33} decay with corresponding rates γ_{ij} ($i = 3, 4$ and $j = 1, 2$) into the ground state populations ρ_{22}, ρ_{11}.
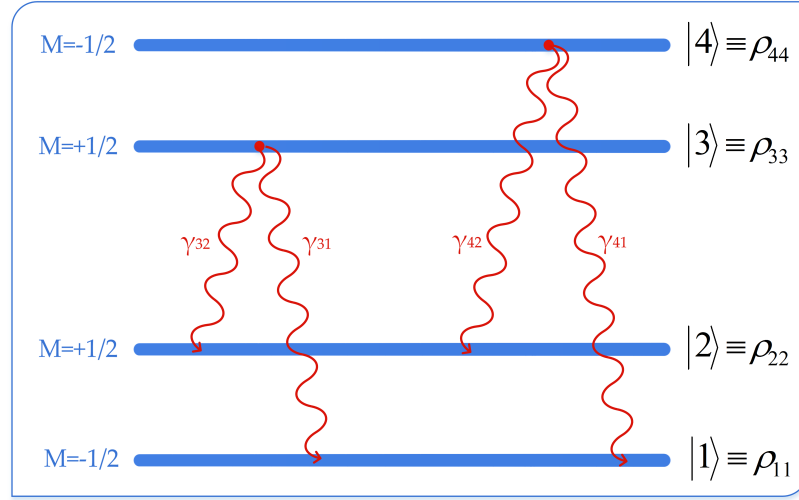
Figure 4: Schematic depiction of the decay processes in the four-level system of our ansatz. The excited state populations ρ_{44}, ρ_{33} decay into the ground states.

The coherences $\rho_{ij}, i \neq j$, represent the magnitude of well-defined, constant phase relation, i.e. coherence, caused by the external energy field (the laser) between state $|i\rangle$ and all states it has coherence with. They decay with rates $\sum_j \gamma_{ij}/2$ [41].

Following these arguments leads to a $\hat{\rho}_s$ in matrix form of [23]:

$$
\begin{pmatrix}
\gamma_{31}\rho_{33} + \gamma_{41}\rho_{44} & 0 & -1/2(\gamma_{31} + \gamma_{32})\rho_{13} & -1/2(\gamma_{41} + \gamma_{42})\rho_{14} \\
0 & \gamma_{32}\rho_{33} + \gamma_{42}\rho_{44} & -1/2(\gamma_{31} + \gamma_{32})\rho_{23} & -1/2(\gamma_{41} + \gamma_{42})\rho_{24} \\
-1/2(\gamma_{31} + \gamma_{32})\rho_{31} & -1/2(\gamma_{31} + \gamma_{32})\rho_{32} & (\gamma_{32} + \gamma_{31})\rho_{33} & -1/2(\gamma_{31} + \gamma_{32})\rho_{34} \\
-1/2(\gamma_{41} + \gamma_{42})\rho_{41} & -1/2(\gamma_{41} + \gamma_{42})\rho_{42} & -1/2(\gamma_{41} + \gamma_{42})\rho_{43} & (\gamma_{42} + \gamma_{41})\rho_{44}
\end{pmatrix}.
$$

Using the Clebsch-Gordan coefficients $C(j_1 j_2 J, M_1 M_2 M)$ and the spin (I_g, I_e) as well as magnetic quantum numbers (M_g, M_e), we can relate the partial

decay rates to the total decay rate Γ of the system [27]:

$$\gamma_{ji} = \frac{2I_e + 1}{2L + 1} C^2(I_g I_e L, M_g - M_e M) \, \Gamma =: \frac{2I_e + 1}{2L + 1} C_{ij}^2 \, \Gamma, \qquad (2.45)$$

where we let i set the quantum numbers of the ground state g, and j the ones of the excited state e. Thereby, C_{ij} equals $C(I_i I_j L, M_i - M_j M)$ with the state $|j\rangle$ defined as $|I_e M_e\rangle$ and $|i\rangle$ as $|I_g M_g\rangle$. L depicts the multipolarity of the decay photon.

We will denote $\frac{2I_e+1}{2L+1} =: \beta$ in the subsequent MBE (Eq. (2.46)).

2.7 MAXWELL-BLOCH EQUATIONS

As we have done all the preparations in the section before, we are just left with combining our previous findings.

In summary: the MBE consider not only the system dynamics arising directly from QM via Eq. (2.6), but also include the spontaneous decay (section 2.6) as well as the response the incident radiation generates in the sample.

To obtain the form below, we exploit that $C(I_e I_g 1; M_e - M_g M) = C(I_g I_e 1; M_g - M_e M)$ with the respective ground (g) and excited (e) state level angular momenta of our system, and honoured the scaling factor $(-1)^{I_g - M_g}$ from equation (2.34).

$$\partial_t \rho_{11} = \beta\Gamma(C_{31}^2\rho_{33} + C_{41}^2\rho_{44}) - {}^{i}\!/\!{}_2 C_{41}(\Omega^*\rho_{41} - \Omega\rho_{14}),$$

$$\partial_t \rho_{21} = -2i\Delta_g\rho_{21} + {}^{i}\!/\!{}_2(C_{32}\Omega^*\rho_{31} + C_{41}\Omega\rho_{24}),$$

$$\partial_t \rho_{31} = {}^{-\beta\Gamma}\!/\!{}_2(C_{31}^2\rho_{31} + C_{31}^2\rho_{31}) - i(\Delta_g - \Delta_e + \Delta)\rho_{31}$$
$$\qquad + {}^{i}\!/\!{}_2\Omega(C_{32}\rho_{21} + C_{41}\rho_{34}),$$

$$\partial_t \rho_{41} = {}^{-\beta\Gamma}\!/\!{}_2(C_{41}^2 + C_{42}^2)\rho_{41} - i(\Delta_e + \Delta_g + \Delta)\rho_{41} - {}^{i}\!/\!{}_2 C_{41}\Omega(\rho_{11} - \rho_{44}),$$

$$\partial_t \rho_{22} = \beta\Gamma(C_{32}^2\rho_{33} + C_{42}^2\rho_{44}) + {}^{i}\!/\!{}_2 C_{32}(\Omega^*\rho_{32} - \Omega\rho_{23}),$$

$$\partial_t \rho_{32} = {}^{-\beta\Gamma}\!/\!{}_2(C_{31}^2 + C_{42}^2)\rho_{32} - i(\Delta - (\Delta_g + \Delta_e))\rho_{32} + {}^{i}\!/\!{}_2 C_{32}\Omega(\rho_{22} - \rho_{33}),$$

$$\partial_t \rho_{42} = {}^{-\beta\Gamma}\!/\!{}_2(C_{41}^2 + C_{42}^2)\rho_{42} - i(\Delta_e - \Delta_g + \Delta)\rho_{42} - {}^{i}\!/\!{}_2\Omega(C_{41}\rho_{12} + C_{32}\rho_{43}),$$

$$\partial_t \rho_{33} = -\beta\Gamma(C_{32}^2 + C_{31}^2)\rho_{33} + {}^{i}\!/\!{}_2 C_{32}(\Omega\rho_{23} - \Omega^*\rho_{32}),$$

$$\partial_t \rho_{43} = {}^{-\beta\Gamma}\!/\!{}_2(C_{41}^2 + C_{42}^2)\rho_{43} - i2\Delta_e\rho_{43} - {}^{i}\!/\!{}_2(C_{41}\Omega\rho_{13} + C_{32}\Omega^*\rho_{42}),$$

$$\partial_t \rho_{44} = -\beta\Gamma(C_{41}^2 + C_{42}^2)\rho_{44} - {}^{i}\!/\!{}_2 C_{41}(\Omega\rho_{14} - \Omega^*\rho_{41}),$$

$$\left(\partial_z + \frac{1}{c}\partial_t\right)\Omega(z,t) = i\eta\left(C_{32}\cdot\rho_{32}(z,t) - C_{41}\cdot\rho_{41}(z,t)\right).$$

$$(2.46)$$

By solving the above equations for the matrix element $\Omega(z,t) \sim E_k(z,t)$ we expect to find the amplitude of our radiation as and after it passes through the grid of nuclei.

As initial conditions we choose:

- The initial populations are 50% in each of the two ground states ρ_{11}, ρ_{22} while all other ρ_{ij} are set to zero.

- The incident radiation is given by an ultrashort pulse, which we smooth, for computational purposes, to a Gaussian of width $\tau = 1ns$ (in intensity) that is reaching the sample at 10τ: $\Omega(0,t) = \exp\left(-\frac{(t-10\tau)^2}{\tau}\right)$.

Part II

NUMERICAL RESULTS

This part presents the numerical results of this thesis. It begins
with the reproduction of two fundamental effects utilized and then
shows the new findings. It concludes with a summary and outlook.

<div style="text-align: right; font-size: 3em;">3</div>

COHERENT PHOTON STORAGE AND PHASE MODULATION

In figure 2 on page 9, we have specified a setup, in which an x-ray pulse impinges on a probe of ^{57}Fe Mößbauer nuclei. The pulse is tuned to the 14.4 *keV* M1 transition and due to its sufficient width, it addresses all six transitions between the six magnetic sublevels. However, only the two $\Delta M = 0$ transitions are driven as a consequence of the combination of the linear pulse polarization and the choice of the quantization axis (cf. sec. 1.2 on page 8). The external magnetic field determining that axis permeates the sample and can be switched on and off as well as inverted.

This chapter is dedicated to explaining the fundamental behaviour of the forward scattered electric field in the time spectrum. Furthermore, we reproduce the results of Ref. [1], which presented the coherent photon storage and phase modulation via magnetic switching.

3.1 RESONANTLY SCATTERED FIELD AMPLITUDE

The dynamics of our incident electric field, as it passes through the sample, are given by $\Omega(z, t)$ from sec. 2.7. Being interested in the forward scattered intensity detectable behind the sample, we view Ω's time evolution at the end $z = L$ of the medium.

Figure 5.**i)** illustrates the unperturbed time spectrum dynamics with and without Zeeman splitting, as they are explained in the following. The plot shows our results from numerically solving the MBE derived in chapter 2 for optical thicknessess $\xi = 60$, $\xi/2 = 30$, magnetic fields $B = 34.4 \, T$, $B = 0 \, T$ and a time range of 600 *ns*.

Dynamics without Zeeman Splitting

In the absence of any external magnetic field, only the two degenerate energy levels of each the ground and first excited state of our specified nucleus system ^{57}Fe are within relevant scope of the laser. The laser can therefore drive only this one transition between the two levels.

Reflecting the driving of a single transition, one could expect to see only a natural, spontaneous decay $\Omega \sim e^{-\Gamma t}$ of the unstable excited state.

However: our equations contain the optical thickness ξ, defined in sec. 2.5 on page 27, which is proportional to the number of nuclei in the sample. As we choose a sufficiently large thickness, the resonant photon gets reabsorbed and emitted by several nuclei before leaving the medium. This is referred to as the "Multiple Scattering Effect".

The recoilless absorption of the photon leads to a re-emission within the line width of the just driven transition (see sec. 1.2). Thereby, its remaining energy is sufficient to again drive this same transition. As this effect may sequentially happen with multiple nuclei before the beam leaves the sample, the scattered radiation arises from a superposition of multiple scattering processes. The interference of this superposition constructs a pattern called Dynamical Beat, which is shown in figure 5.

An analytic solution of this Dynamical Beat via the first Bessel function J_1 is provided as [23]

$$|\Omega_0(L,t)|^2 \sim \left(\frac{\xi}{\sqrt{\xi\Gamma t}} J_1\left(2\sqrt{\xi\Gamma t}\right) \right)^2 e^{-\Gamma t},$$

which is proportional to $\frac{1}{\xi\Gamma}$ at short times. We obtain the same dynamics by numerically solving our system of PDE (2.46) for $B = 0$.

As ξ grows, so does the frequency of oscillations in Ω_0. Nonetheless, the overall lifetime of these dynamics is naturally limited by the spontaneous decay $e^{-\Gamma t}$.

Dynamics with Zeeman Splitting

Immersing the nuclei in a non-vanishing magnetic field removes the degeneracy from the energy levels, so that the ground state splits into two and the
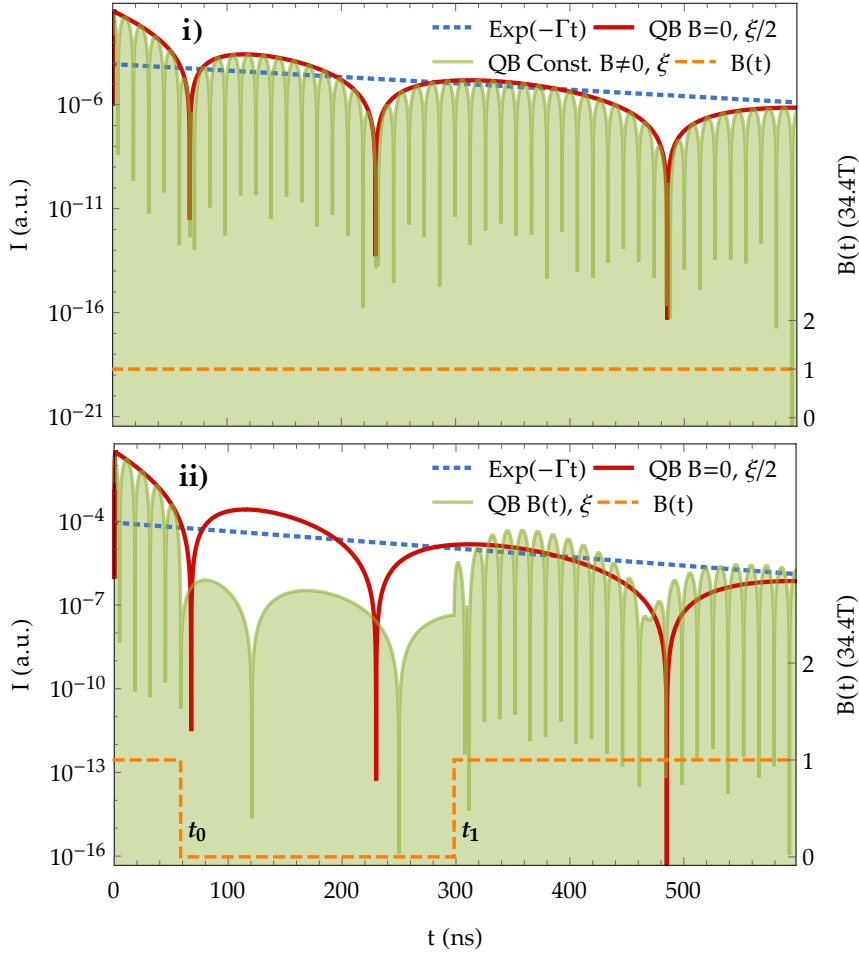
Figure 5: **i)** Quantum beat of the $\Delta M = 0$ transitions (green line) for a constant magnetic field $B = 34.4\ T$ and optical thickness $\xi = 60$. The corresponding envelope (red line) is given by the Dynamical Beat calculated for $B = 0$ and $\xi = 30$. See text for further explanations. **ii)** Coherent storage effect between t_0 and t_1, the storage time can be arbitraly chosen and was set to $\tau_{off} = 240 ns$. In both plots we used an optical thickness of $\xi = 60$ and $\xi/2 = 30$ respectively.

excited into four distinctive sublevels. We end up in our known four-level scheme (figure 3 on page 17) by choosing a geometry in which the selection rule $\Delta M = 0$ forbids all but two transitions. Now, both of these transitions are driven by the laser.

Again, rooted in the delocalized nature of the Nuclear Exciton, we observe an interference. Though, this time, not only of multiple scatterings, but additionally of the two transitions $\hbar\omega_{32}$ and $\hbar\omega_{41}$ via which the x-ray photon scatters. This holds as long as we only detect the total photon count behind the sample and do not resolve the detected radiation in the energy spectrum. The carrier oscillations are governed by the beat of ω_{32} and ω_{41}, called Quantum Beat

(QB).

The envelope dynamics is still given by the multiple scatterings and therefore the Bessel-function. We note here that for the same optical thickness ξ, the Dynamical Beat (for the case of $B = 0$) does not represent the QB envelope. We can see in figure 5 that the optical thickness is ξ for the QB and only $\xi/2$ for the Dynamical Beat that illustrates its envelope. The pace of the Dynamical Beat (for $B = 0$) is connected to the number of nuclei in the same ground state via the multiple scattering. In case of $B = 0$, the number of resonant nuclei is assumed to be N. Since the ground state splits into two Zeeman levels for $B \neq 0$ and the populations distribute equally at room temperature, only $N/2$ nuclei contribute to the resonant scattering for each of the two $\Delta M = 0$ transitions. As a result, at the same ξ, the QB envelope oscillates slower than the Dynamical Beat corresponding to the pure two-level system in the absence of any magnetic field.

Since ξ is proportional to the number of nuclei (see the definition on page 27), we compensate for this effect by halving ξ for the envelope plot.

The angular beat frequency ω_{QB} is

$$\frac{\omega_{41} - \omega_{32}}{2} = \frac{(\omega_4 - \omega_1) - (\omega_3 - \omega_2)}{2}$$
$$\overset{\text{fig. 3}}{=} \Delta_e + \Delta_g = \Delta_{B_0} = \omega_{QB}. \tag{3.1}$$

We may infer from the amplitude dynamics that two beats make up one period T_{QB}. Hence, we witness one lope every $T_{QB}/2 = \pi/\Delta_{B_0}$ time steps. In case the envelope (Dynamical Beat for half ξ) and beat minima would not coincide, the QB regularity is broken so that the lopes can be squeezed into the envelope and the minima do coincide. This squeezing effect is visible in figure 5.**i)** around the dynamical beat minima.

With the Zeeman splittings being proportional to the magnetic field strength, we can narrow the QB pattern by increasing B. Throughout this thesis, we only consider $\frac{1}{\Gamma} > \frac{1}{\Delta_{B_0}}$ to prevent too complex beat structures.
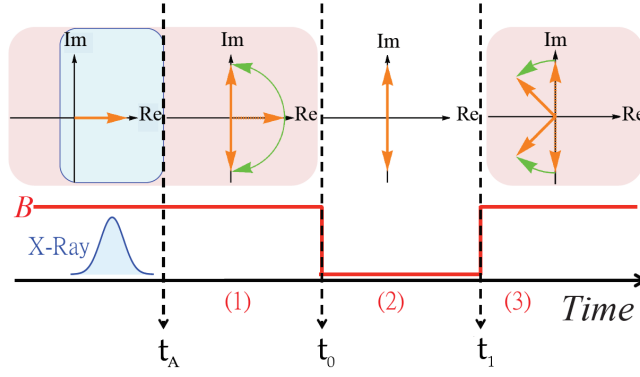
Figure 6: Visualization of the rotating current matrix elements. The x-ray reaches the target at t_A. During time interval (2), the magnetic field is switched off and the current matrix elements (see Eq. 2.41) are frozen on the imaginary axis. Image courtesy of Wen-Te Liao [23].

3.2 STORAGE EFFECT

The idea of storing a photon emerges from manipulating the Zeeman splitting by applying a time-dependent B-field. In particular, the switching between a vanishing and a non-vanishing B at specific times is the essence of the storage effect.

For the time we apply a non-vanishing field, we expect to observe the QB. Our results in figure 5.**ii)** show: at the instance of turning B to zero, the pattern switches into the Dynamical Beat. The time t_0 controls the suppression of the Dynamical Beat after switching (green) versus the unperturbed one (red). Hence, the highest signal suppression is obtained by switching in a QB minimum.

Switching the magnetic field off at time t_0 ceases the QB and stores the photon. By turning B back on at t_1 the photon is released and we can observe the coherent storage effect: the QB continues its dynamics exactly at the point it was ceased before. As we will see in the next chapter, this storage process is repeatable and can be performed several times.

An explanation [1] for the storage effect is given by figure 6. As long the magnetic field is present, the two transition current matrix elements (Eq. (2.41)) rotate in the complex plane with $\omega_{QB} = \Delta_B$ relative to each other. They both start parallel on the real axis, which is why they destructively superimpose on

the imaginary axis after half a period $T_{QB}/2$ (which is the instance of a QB minimum).

If we switch off the field at this instant, the rotation freezes and the destructive interference of the currents holds for as long as the field stays off.

This scheme conserves photon energy, polarization and phase [23] of the co-herently stored photons. However, we can see that storage time, just as the unperturbed signal lifetime, is still limited by incoherent natural decay processes draining energy from the system.

3.3 PHASE MODULATION

Rotating the magnetic field by $\pi\ rad$, i.e. inverting its value, causes the reversed time effect [1] in which the nuclear currents evolve backwards in time. This induces a shift in the photon phase of $\pi\ rad$ (fig. 7), which is independent of an eventual photon storage period.

We demonstrate this phase shift effect in figure 7 by implementing an according magnetic field inversion at time t_s into the MBE (2.46). We set $\zeta = 60$ and choose the storage time to be zero.

Having demonstrated the fundamental effects of switching the B-field without an alteration of the (sign-insensitive) direction of the quantization axis, we proceed to investigating the features of more complex magnetic switching sequences in the next chapter.
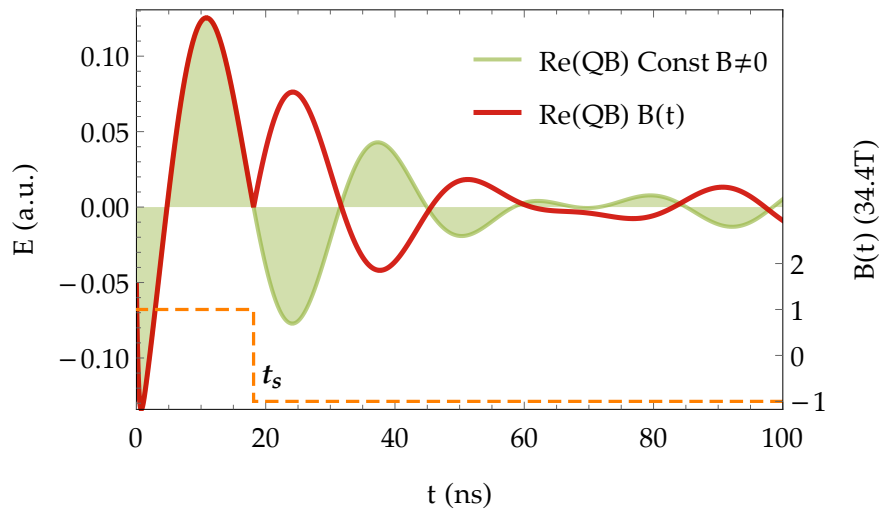
Figure 7: Time evolution of the real part of the scattered radiaton at position L. The magnetic field is rotated by π rad at time t_s. This field inversion induces a phase shift in the electric field $E(L, t)$ wave of π rad.

4

GENERATION AND CONTROL OF FREQUENCY COMBS

This chapter is dedicated to investigating the effects of periodic magnetic switchings on the NFS spectrum. The previously introduced photon storage and phase modulation (chapter 3) are repetitively performed in periodic patterns. Not only the time, yet also the frequency spectrum is analysed. We find that the periodicity in the switching generates frequency combs with equidistant peaks in the forward scattered radiation. The combs' shape can be controlled via the periodic properties of the magnetic switching.

To document the method used for viewing the frequency domain, we initially introduce the Fourier transform utilized to switch into frequency space. For verification of the implemented transform, we compare our results for the unperturbed QB with the theoretical expectations.

Consecutively, we systematically simulate time and frequency domains under the imposition of periodic magnetic field switchings. We proceed by analysing the qualitative behaviour of the combs.

Finally, we present a semi-analytical description of the scattered spectra. We construct an integrable function that, in first order, simulates the time spectrum. This permits us to find the analytic Fourier transform qualitatively leading to the observed frequency combs.

Throughout this chapter, we set the optical thickness ξ sufficiently low to suppress multiple scattering processes for the duration of our scope. In good approximation, the QB envelope is then given by the exponential decay $e^{-\Gamma t}$ [29]. But more importantly, the beat lopes becomes equidistant, so that it would become easier to experimentally implement a periodic switching sequence aligned with the beat.

4.1 FOURIER TRANSFORM

As mentioned in the chapters before (2, 3), we obtain our forward scattered light dynamics $\Omega(z,t)$ from numerically solving a set of PDE, the Maxwell-Bloch Equations (MBE, Eq. (2.46) on page 30). Up to this point, there is no analytic solution of these equations available. Therefore, we can only resort to the, yet very exact, numerical approximation and have to turn to a discrete Fourier transform instead of an analytic one.

Given a finite complex series $u = (u(1),\ldots,u(N))$ of N sampled function values, the transform is implemented in Wolfram Mathematica as

$$\mathcal{F}[u] =: \tilde{u}(\omega) = \frac{1}{\sqrt{N}} \sum_{t=1}^{N} u(t) e^{-2\pi i(t-1)(\omega-1)/N}. \tag{4.1}$$

As the indices run form 1 to N, we cannot implement negative frequencies directly.

The solution comes from assuming the input u to be an N-periodic sequence given in equidistant entries (of distance Δt). The transform holds the same periodicity. Therefore, higher, next period frequencies correspond to negative frequencies:

$$\tilde{u}(\omega + N) = \tilde{u}(\omega), \text{ so that}$$

$$\tilde{u}(N-1) = \tilde{u}(-1). \tag{4.2}$$

With the zero frequency term given at the first position ($\tilde{u}(1)$), we shift our data points by $\frac{N-1}{2}$ entries to center the spectrum around zero.

By transform default, the width of the frequency range (f_{range}) is given as an integer $f_{range} = N$ with the equidistant steps Δf on the frequency scale corresponding to $\Delta f = 1$.

To obtain viable results, we normalize f_{range} so that it corresponds to the period T in the time spectrum, $f_{range} \overset{!}{=} \frac{2\pi}{\Delta t}$, and so that the unit steps Δf are meaningful: $\Delta f = \frac{f_{range}}{N} = \frac{2\pi}{N\Delta t}$. Now, a step in the frequency domain has a fixed translation into an actual frequency unit that is directly comparable with the time spectrum unit.

For normalization of the y-axis, we scale $\tilde{u}$ by a factor $\frac{T}{\sqrt{2\pi N}}$ so that it matches the continuous Fourier transform.

In figure 8.**i)** we simulate the time spectrum for the parameters $\xi = 0.5$ and
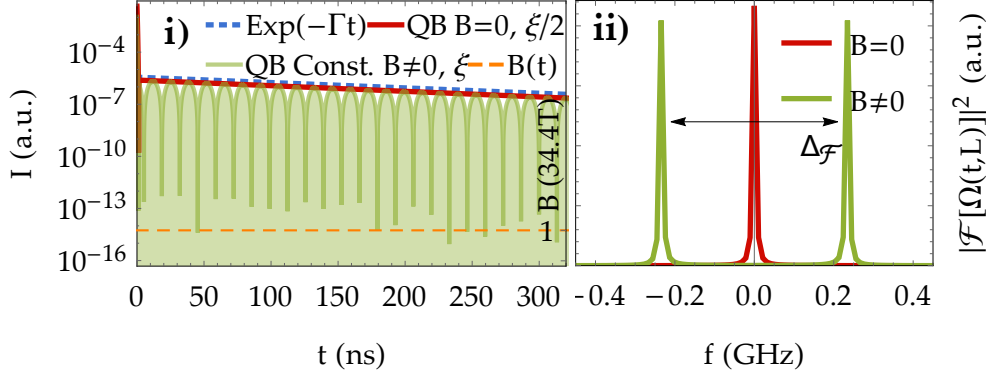
Figure 8: **i)** Time spectrum for a constant magnetic field and optical thickness $\xi = 0.5$. The Dynamical Beat agrees with the natural exponential decay, while the QB shows a decaying, absolute-cosine-like oscillation. **ii)** The Dynamical Beat consists of the ω_0-frequency, the QB incorporates the two transition frequencies ω_{41} and ω_{32}.

a constant magnetic field $B = 34.4\ T$. The MBE are numerically solved over a time range of $600\ ns$, which is used for all following MBE solutions as well. The length of the time range consigned to the transform determines its resolution, the larger the range the higher the resolution in the frequency spectrum. Part **ii)** of figure 8 shows the results for the Dynamical and the unperturbed Quantum Beat in the frequency domain. It is crucial to bear in mind, that we see frequencies relative to $\omega_L \overset{\Delta=0}{=} \omega_0$ due to our transformation into the rotating frame (see chapter 2).

As expected, for $B = 0$ only the principal transition frequency ω_0 occurs. With a constant magnetic field $B \neq 0$, we observe the frequencies of the two Zeeman-split transitions ω_{41}, ω_{32}. We may verify our transform by reading out their distance $\omega_{41} - \omega_{32} =: \Delta_{\mathcal{F}}$.

Since the Zeeman-splitting $\Delta_{B_0} = \Delta_g + \Delta_e$ (cf. fig. 3 on page 17) equals the QB frequency ω_{QB}, we expect for the marked peak separation in plot **ii)**:
$$\frac{\Delta_{\mathcal{F}}}{2} = \omega_{QB} = \Delta_{B_0}.$$

With $\Delta_{\mathcal{F}} \approx 0.46915 GHz$ and $2\Delta_{B_0} \approx 0.4612 GHz$, we incur a divergence of 0.006%. Considering the limited resolution, we view our transform as verified. Note that, to easily exempt the direct response from the incident pulse itself and prompt scattering processes close to $t = 0$ other than the resonant NFS, we only transform from the first QB minimum to the last in the time range over which we solved the MBE. These minima are at 5 ns and 595 ns respectively

in the time range shown in the plots. We presume that the key characteristics of the following observations remain when instead transforming the full spectrum with the initial light pulse cut off as it is shown in the plots. However, since that would include the signal of other scattering processes with no or different periodicity, the frequency spectra would gain noise.

4.2 QUALITATIVE ANALYSIS

We now proceed to simulate the spectra when the probe is immersed in periodically switched magnetic fields.

We have implemented two main kinds of switching sequences:

- a non-inverting one (NInv), alternating between the magnetic field at full strength and zero, and

- an inverting one (Inv), switching between B at full strength, zero and B at full strength rotated by $180°$.

The critical parameters of these sequences are the time intervals of the different stages. We denote the storage time, during which the field is not present, as τ_{off}. As seen in the preceding chapter 3, this period can be chosen without constraints.

The B-field on-time we refer to as τ_{on}. As to our aim of achieving maximum signal suppression, τ_{on} is restricted to multiples of $\frac{T_{QB}}{2}$.

Non-Inverting Sequence

We implement the non-inverting sequence into the MBE and solve them for different storage times in combination with a fixed τ_{on} of half a QB period. As before in figure 8, ξ is set to 0.5 and B_0 to 34.4 T. The numerical results in figure 9 reveal: under the influence of the periodic switching, the two peaks seen before in figure 8 split into several equidistant spikes in the frequency domain. Thus, a frequency comb is formed.

To check the equidistance, we evaluated the peak positions. They are drawn as black-dotted lines into the graphs. The average peak distance has a standard deviation of 0% up to circa 5% for different combs. We attribute this deviation to the limited resolution and thereby limited precision of the peak positions.
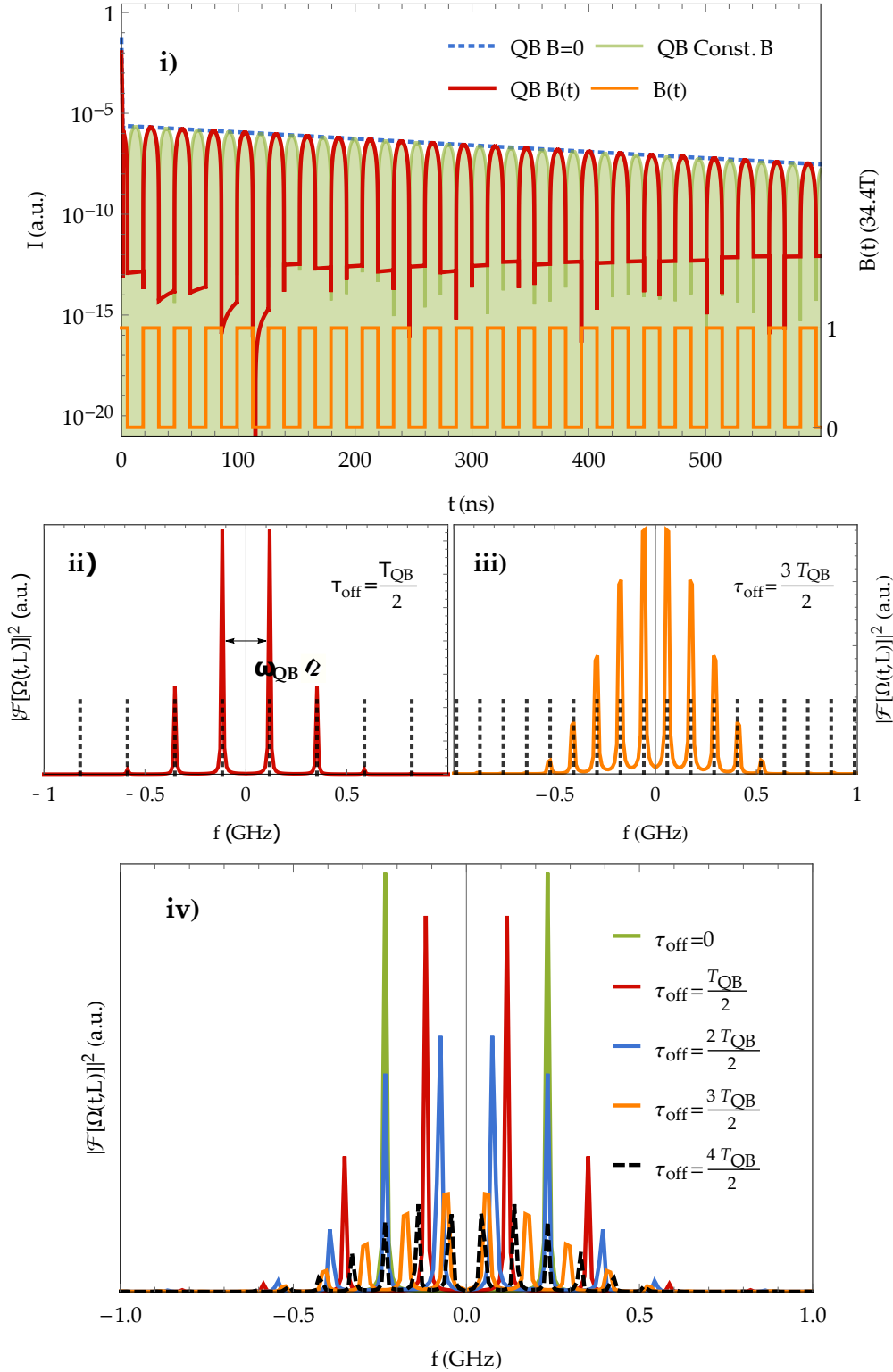
Figure 9: **i)** Time spectrum for $\tau_{on} = \tau_{off} = \frac{T_{QB}}{2}$ (red). **ii)** and **iii)** Single frequency spectra for different τ_{off} at fixed $\tau_{on} = \frac{T_{QB}}{2}$. The black dotted lines mark the equidistant peak positions. **iv)** Multiple frequency spectra for different τ_{off} in comparison. With a growing storage time, the peaks multiply and narrow towards ω_0 while the overall intensity drops. The spectra of $\tau_{off} = 0$ and $\frac{T_{QB}}{2}$ have been manually downscaled by 80% and 50% respectively to allow for a proper display.

This is supported by the fact, that there is no systematic or biased divergence acknowledgeable.

In figure 9, the signals for varying τ_{off} and τ_{on} are displayed. From this and further not depicted simulations, we conclude the following:

- At constant τ_{on}, the separation of the peaks in the frequency combs narrows with growing τ_{off}. The peaks drift towards the origin (ω_0) and new peaks gain intensity at the border regions (towards $\pm 1 GHz$). The two frequencies ω_{41}, ω_{32} appear in the combs for every τ_{off} set to an even multiple of $\frac{T_{QB}}{2}$. Due to maximum signal suppression during τ_{off}, the overall intensity declines as we increase this parameter.

- For a constant τ_{off} but a varying τ_{on}, we may lose the equidistant peak separation: if τ_{on} is sufficiently large to allow only for a very few switchings, we have an almost unaltered time spectrum in comparison to the constant magnetic field. In consequence, we see almost just the two constant-field beat frequencies. But, as becomes evident for reducing τ_{on}, the two frequency peaks at ω_{41}, ω_{32} split into two spikes. Their separation grows for a declining τ_{on}. No additional high-intensity peaks emerge. We have done simulations with τ_{on} in a range from $\frac{T_{QB}}{2}$ to the point at which the field is only switched once, $\frac{45 T_{QB}}{2}$.

- We tested keeping the relation between switching times constant, $\tau_{on} = \tau_{off} =: \tau_S$. Increasing τ_S produces a mixture of the previous two scenarios. The peaks are narrowing, yet not towards ω_0, but ω_{41} and ω_{32}. At uneven multiples of $\frac{T_{QB}}{2}$, we see the two highest peaks adjacent to ω_{41}, ω_{32}. At even multiples they appear exactly at those two points.

We conclude that suitable frequency combs may only be generated for $\tau_{on} = \frac{T_{QB}}{2}$.

Inverting Sequence

This sequence differs from the NInv one in that the orientation of the *B*-field is inverted after every off-switching. This 180° rotation causes a π-phase shift in the time domain (see fig. 7 on page 39). As this phase shift is equivalent to flipping every second beat lope with respect to the time axis, the intensity time spectrum remains the same as before.
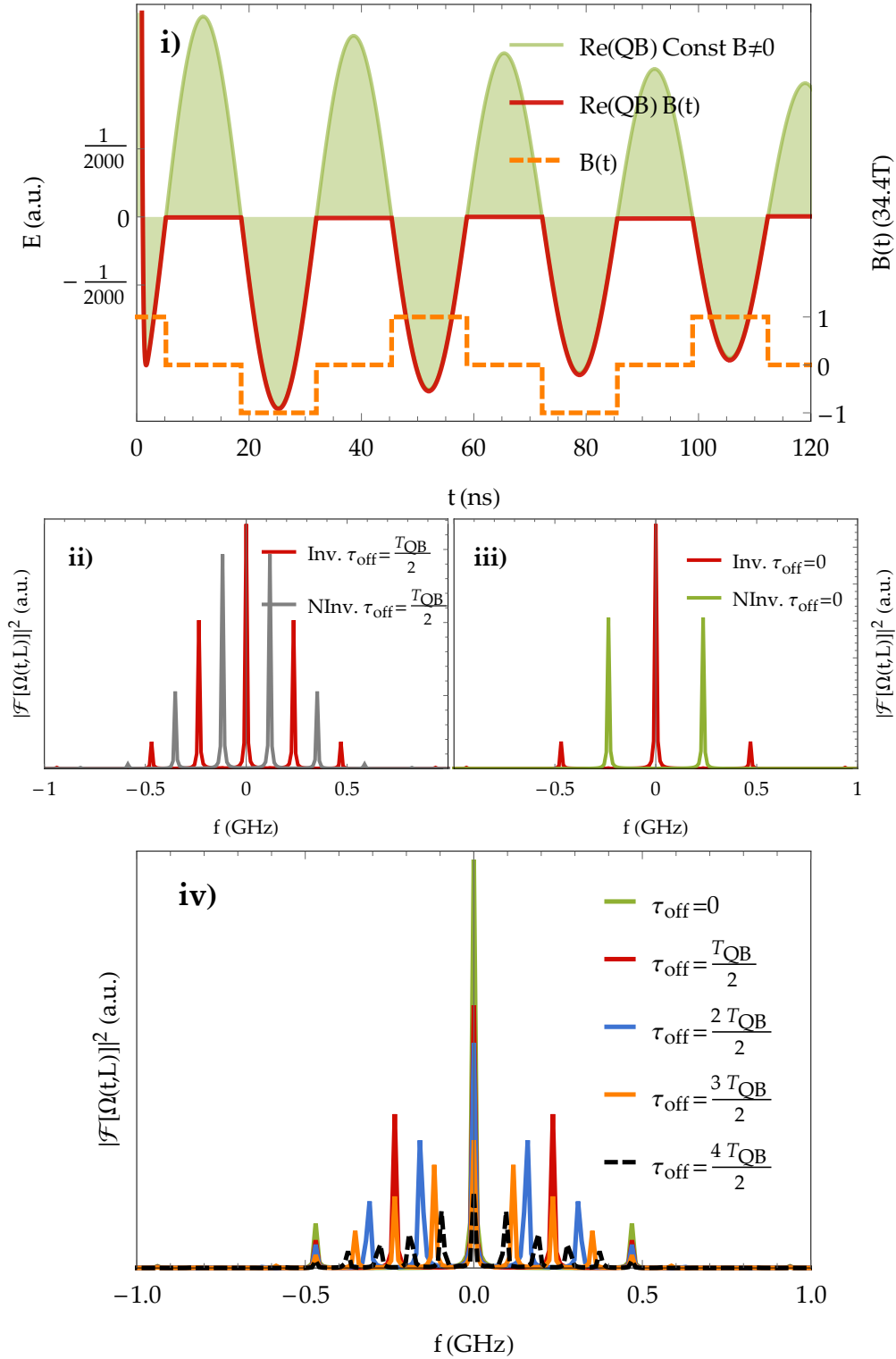
Figure 10: **i)** *E*-field amplitude time spectrum for the inverting sequence at $\tau_{on} = \tau_{off}$ set to half the QB period. The field direction is inverted after every storage. **ii)** and **iii)** Frequency combs for $\tau_{off} = \frac{T_{QB}}{2}$ and $= 0$ respectively in comparison with the inverting and non-inverting switching. **iv)** Frequency combs of the inv. sequence for several different storage times. The spectra of $\tau_{off} = 0$ and $\frac{T_{QB}}{2}$ have been manually downscaled by 80% and 50% respectively to allow for a proper display.

In figure 10, we present the implementation of the Inv sequence for several storage times τ_{off} ranging from 0 to $^{4T_{QB}}/2$. We give an example for the scattered field (real and imaginary part), that demonstrates the functionality of the switching sequence. Furthermore, we compare the Fourier spectra for the Inv and NInv series for two exemplary storage times and show how controlling the storage time influences the comb shape.

Interestingly, the Fourier spectra of the Inv switching differ significantly from the NInv case. The frequency combs generated with the inverting sequence all have their highest peak at ω_0.

For varying τ_{off} (shown in fig. 10.**iv)**) and fixed τ_{on} as done for the NInv sequence above, these combs show an analogous behaviour. The ω_0-peak vanishes for even n in $\tau_{on} \geq \frac{nT_{QB}}{2}$.

As we can infer from figure 10.**ii)**, the Inv combs are shifted relative to the NInv combs by half their equidistant separation.

Another remarkable effect comes up in figure 10.**iii)** when setting the storage time to zero. Then, the magnetic field is constantly and immediately inverted after τ_{on}. The phase shift results in an *E*-field time spectrum that resembles a flipped absolute cosine. The intensity spectrum is the same as with a constant *B*-field.

However, in contrast to the QB in the case of a constant field, we do not see the two transition frequencies. Instead, the frequency spectrum shows a large intensity peak at ω_0 and two smaller ones each separated by Δ_{B_0} from the origin.

Again, the shift distance between the corresponding ($\tau_{off} = 0$) NInv spectrum peaks (green in fig. 10.**iii)**) and the Inv is half the equidistance, in the plotted case $\frac{\Delta_{B_0}}{2}$.

4.3 ANALYTIC APPROACH

To obtain a better comprehension of the observed phenomena, we attempt a semi-analytic description of the spectra. The term semi-analytic is proposed to convey that we do not analytically solve the underlying differential equations ((2.46) in chapter 2), yet we find an analytic function that models the observed time spectra in first order. This function is then analytically Fourier

transformed and compared to the actual, discrete transform from the preceding sections.

For a constant magnetic field, the time spectrum dynamics can be modelled by a cosine-function combined with an exponential decay term:

$$\Omega(L,t) \approx A \cdot \cos(\omega_{QB}t + \phi)e^{-\frac{\Gamma}{2}(t+t_s)}\Theta(t), \tag{4.3}$$

with a constant amplitude A, phase shift ϕ and shift time t_s to correctly render the starting time of the natural decay. The Heaviside-Theta-function $\Theta(t)$ ensures our spectrum starts at time $t = 0$.

We set $\phi = \frac{-\pi}{2}$ and t_s to the time of the first QB minimum, so that the spectrum is reconstructed from the same origin we have chosen for the discrete transform in the preceding sections.

The general time spectrum for a non-vanishing storage time τ_{off} resembles a series of single, cut-off cosine-beats:

$$\Omega_{single}(L,t) = Ae^{-\frac{\Gamma}{2}t_s}\cos(\omega_0(t - t_1) + \phi)e^{-\frac{\Gamma}{2}t}\Theta(t - t_2)\Theta(t_3 - t). \tag{4.4}$$

The shifting time t_1 controls the starting point of the single cosine beat lope. The other two (t_2 and t_3) set its cut-off positions.

The full spectrum is given by the sequence of these single beats, i.e.:

$$\Omega(L,t) = \sum_{n=0}^{\infty} \Omega_{single}(L,t,n), \tag{4.5}$$

where the n-dependence is introduced in the shifting times $t_i = t_i(n)$, $i = 1, 2, 3$. The explicit form of $t_i(n)$ is stipulated by the switching sequence type; the inverting and non-inverting type only differ in their form of the shifting times.
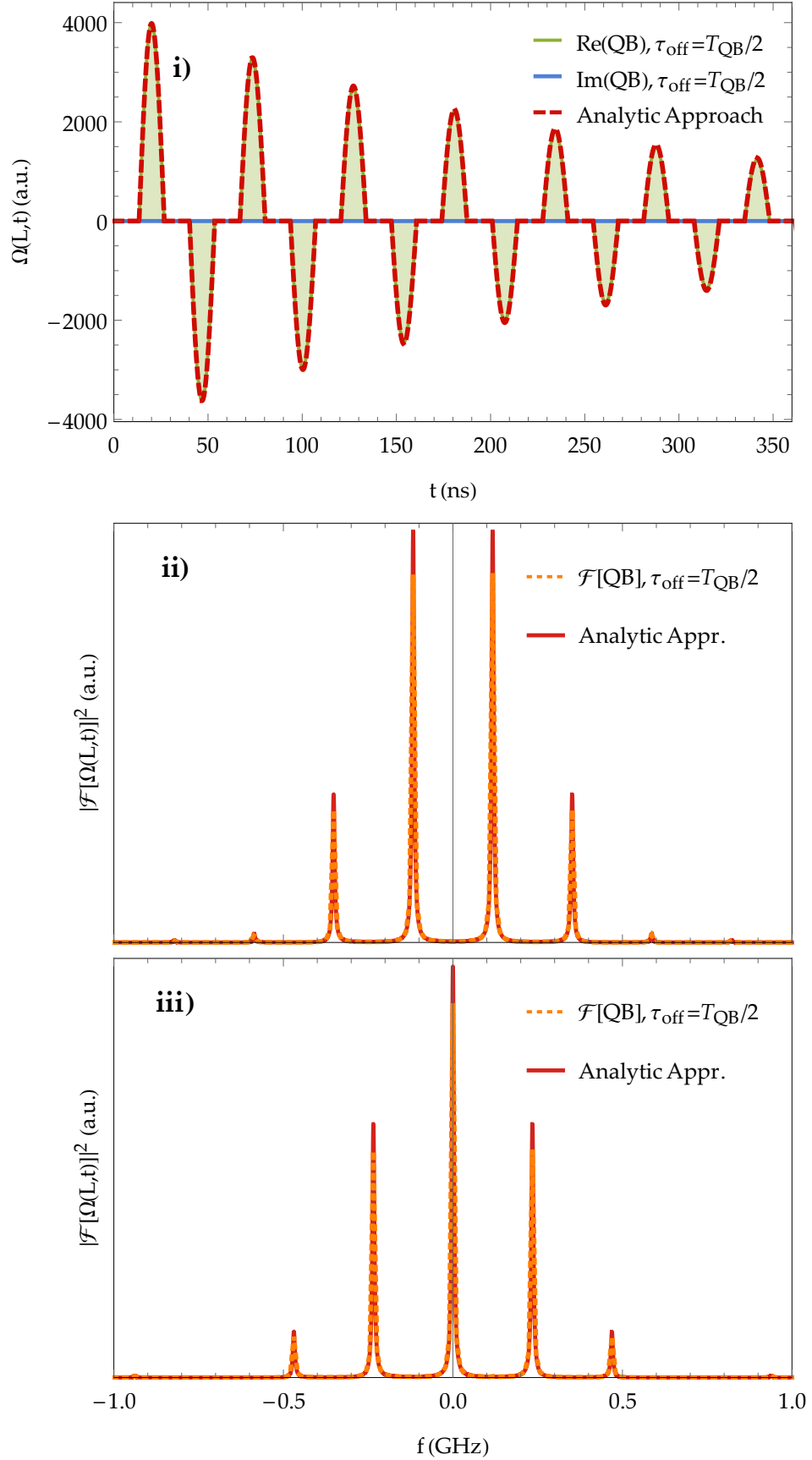
Figure 11: **i)** An example plot of the numerically solved NFS time spectrum from the MBE (green and blue) for the non-inverting sequence at $\tau_{off} = \tau_{on} = T_{QB}/2$ in comparison with the analytic approach defined in section 4.3. The imaginary part of the scattered amplitude is zero. **ii)** and **iii)** Discrete Fourier transforms (orange dashed) compared to the analytic approach (red) for the non-inverting as well as inverting sequence, each at a storage time of half a QB period.

The Fourier transform of the single lope is

$$
\begin{aligned}
\tilde{\Omega}_{single}(L,\omega) &:= \mathcal{F}[\Omega_{single}(L,t)] \\
&= \frac{Ae^{-\frac{\Gamma}{2}t_s}}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \cos(\omega_{QB}(t-t_1)+\phi)e^{-\frac{\Gamma}{2}t} \\
&\quad \times \Theta(t-t_2)\Theta(t_3-t)\cdot e^{i\omega t}\,dt \\
&= \frac{Ae^{-\frac{\Gamma}{2}t_s}}{\sqrt{2\pi}}\left(\frac{e^{i(\phi-\omega_{QB}t_1)}}{i(\omega_{QB}+\omega)-\frac{\Gamma}{2}} \cdot e^{t(i(\omega_{QB}+\omega)-\frac{\Gamma}{2})}\Big|_{t_2}^{t_3} \right. \\
&\quad \left. - \frac{e^{i(\omega_{QB}t_1-\phi)}}{i(\omega_{QB}-\omega)+\frac{\Gamma}{2}} \cdot e^{-t(i(\omega_{QB}-\omega)+\frac{\Gamma}{2})}\Big|_{t_2}^{t_3} \right).
\end{aligned}
\tag{4.6}
$$

As the Fourier transform is a linear functional, we may reduce the transform of the whole spectrum sequence (Eq. (4.5)) to a series of the single transforms:

$$
\begin{aligned}
\tilde{\Omega}(L,\omega) &:= \mathcal{F}\left[\sum_{n=0}^{\infty} \Omega_{single}(L,t,n) \right] \\
&= \sum_{n=0}^{\infty} \mathcal{F}[\Omega_{single}(L,t,n)] = \sum_{n=0}^{\infty} \tilde{\Omega}_{single}(L,\omega,n).
\end{aligned}
\tag{4.7}
$$

Our analytic approach exploits the exponential decay form of the QB envelope for small ξ, it is therefore not valid for a larger optical thickness. It as well takes advantage of the vanishing of the imaginary part of the E-field amplitude, as it is shown in figure 11.**i)**.

In figure 11.**i)** we show an example of the time spectrum modelling for the NInv sequence at a storage time of half a QB period to verify our approach. The optical thickness and time range (for the solution of the MBE, not the plot range) are chosen as in the preceding time spectrum plots (e.g. fig. 9.**i)**). As the plot shows, the analytic time spectrum model agrees, in first order, perfectly with the numerical solution from the MBE (Eq. (2.46)).

Plots **ii)** and **iii)** display the analytic transforms of the Inv and NInv sequences for $\tau_{off} = \frac{T_{QB}}{2}$. The analytic frequency spectrum is in very good agreement with the discretely transformed one from the MBE. The slightly higher intensities in the analytic transform are rooted in the limited resolution of the discrete transform. While the latter is limited to a time spectrum input of the duration for which we solved the MBE (the mentioned time range of approximately 600 ns), the former can resort to a significantly larger time span. Thereby, the

analytic peaks count more beat lopes that constructively contribute to their height.

4.4 FEASIBILITY

Up to this point in this thesis, we have implemented the magnetic field switching as instantaneous. This approach is the simplest to theoretically investigate and understand the behaviour of time and frequency spectra of the scattered photons. But it is not realistic as an infinitely steep switching cannot be implemented in an experimental setup. Therefore, we dedicate this section to test smoothed magnetic switchings with a finite, experimentally more feasible slope.

Initially, we introduce an exponential smoothing of our rectangle pulse function, as shown in figure 12. The smoothing used is described by the following exponential function:

$$
\begin{aligned}
smRect(t, t_0, t_1) := & \left(1 - \frac{1}{1 + \exp\left(\frac{t - t_0}{\epsilon}\right)}\right) \Theta\left(t_0 + \frac{t_1 - t_0}{2} - t\right) \\
& + \frac{1}{1 + \exp\left(\frac{t - t_1}{\epsilon}\right)} \Theta\left(t - \left(t_0 + \frac{t_1 - t_0}{2}\right)\right),
\end{aligned}
\tag{4.8}
$$

with smoothing radius ϵ and $\Theta(t)$ as the Heaviside-Theta function.
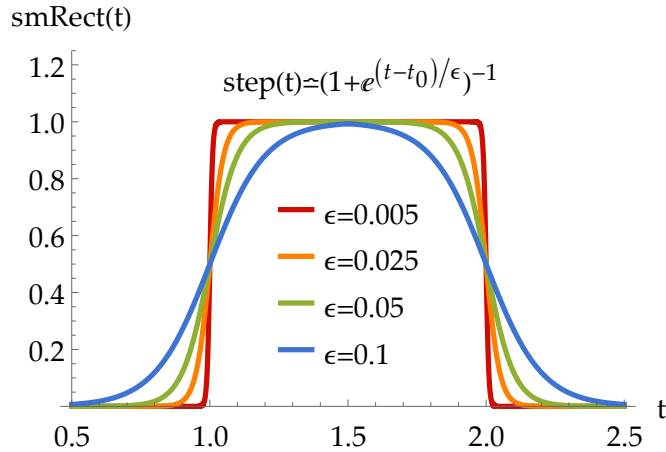
Figure 12: Exponentially smoothed rectangle pulse for step up at $t_0 = 1$ and step down at $t_1 = 2$. The effect of the smoothing radius ϵ depends on the value of those two parameters.

Finally, as an ultimate accommodation for the experimental practicability, we give up one degree of freedom by constraining *B*-field off- and on-time to be the same, $\tau_{off} = \tau_{on}$. This enables the switching sequence to be produced by a sinusoidal signal.

Our simulations with these smoothed sequences show one main result: Up to a certain, significantly large degree of smoothing, the frequency spectra stay basically unaltered with respect to the unsmoothed switching.

As a numerical example, we present figure 13. For the same ξ as in the sections before, figure 13 shows the results of smoothing the switching for the non-inv. sequence at $\tau_{off} = \tau_{on} = \frac{T_{QB}}{2}$. The Fourier transform is performed over the same range as previously, from the first to the last QB minimum (now appr. at 606 *ns*). In part **i)** the switching sequences for different smoothings are compared, so that the degree of smoothing becomes tangible. The smoothing radius for the exponentially smoothed step (orange) was chosen as $\epsilon = 0.5$ *ns*. Plot **ii)** gives an extract of the time spectrum for the differently smoothed QB. A significant deviation of the smoothed versus the abrupt switchings is visible. In part **iii)** we compare the frequency spectra of the various sequences. Except for slight intensity deviations, the main characteristics of the frequency combs are maintained. Even the sinusoidal signal, which is likely to be the simplest to experimentally implement, shows no significant frequency spectrum distortion.

Thus, we conclude that our previous findings bear the realistic opportunity of an experimental implementation.

In terms of the experimental realization, the fundamental constituent is the coherent x-ray source. However, with the advent and anticipated improvement of x-ray free electron lasers (XFEL) [25, 43, 44], another core challenge remaining is the implementation of the ultra-fast magnetic switching.

One possibility of achieving this aim could be via so called snapper capacitors [45] using few-turn coils and a moderate pulse current of approx. 15 kA from the low-inductive high-voltage capacitors [23]. Another interesting prospect is the "Lighthouse Setup" [22, 46] shown in figure 14: Instead of switching the field, the sample is rotated in and out of its reach. A small rotor, which is only a few mm in cross-section, achieves rotational frequencies ω_R of up to 70 kHz. It is expected to be fast enough to rotate the sample out a depth of few μm in
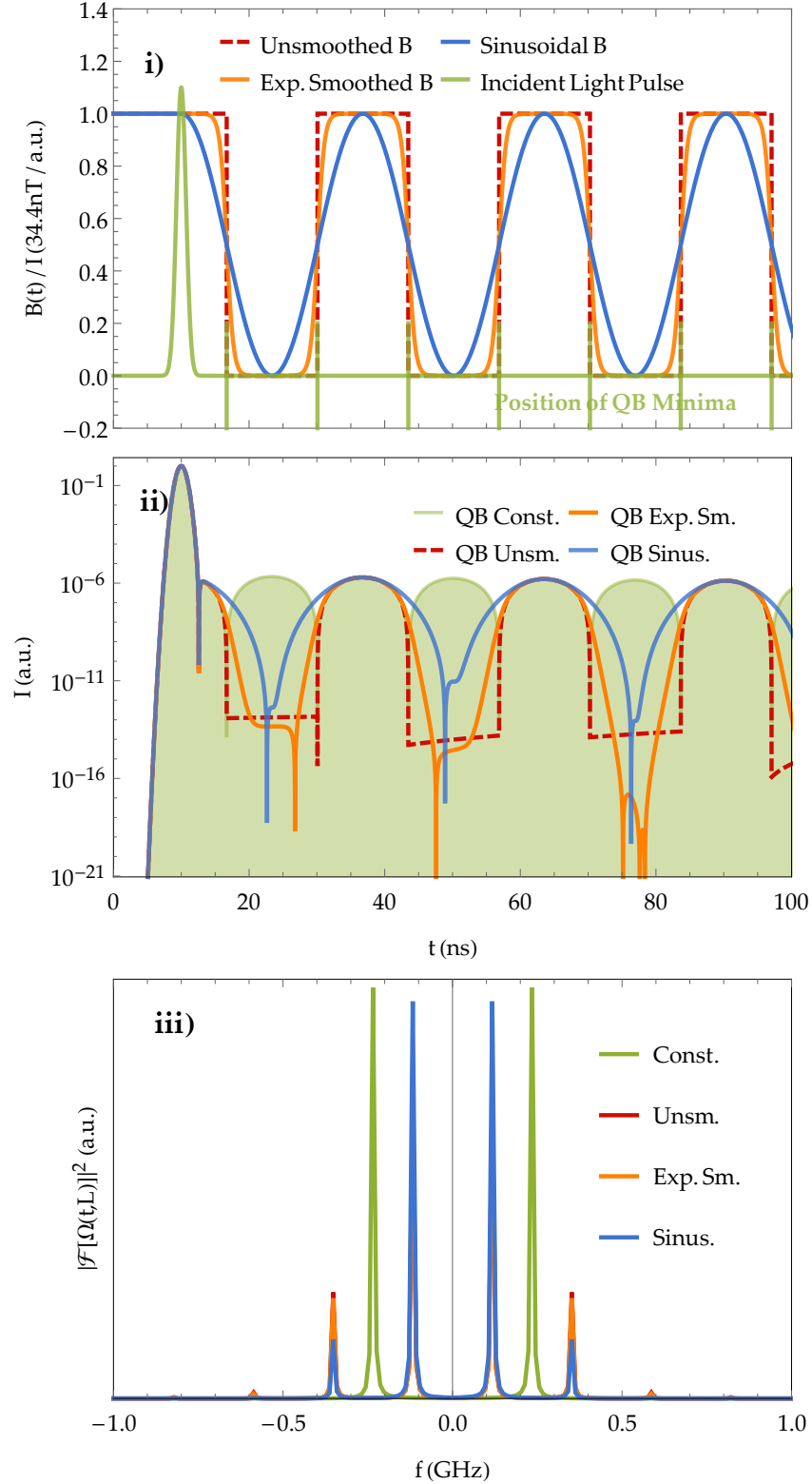
Figure 13: **i)** Abrupt and differently smoothed magnetic switching sequences for $\tau_{on} = \tau_{off} = T_{QB}/2$ in comparison. The smoothing radius for the exponential step (orange) was set to $\epsilon = 0.5$ *ns*. Different than in the previous sections, we display the time spectrum with the origin set before the light pulse arrival at $t = 10$ *ns*. The Fourier transform is performed over the same range as previously, from the first to the last QB minimum (≈ 606 *ns*). **ii)** Extract of the time spectrum for different QB. The origin is chosen the same as in the preceding *B*-field sequences plot. **iii)** Frequency spectra of the various sequences. Except for slight intensity deviations, the main characteristics of the frequency combs are maintained. Even the sinusoidal signal, which is likely to be the simplest to experimentally implement, shows no significant frequency spectrum distortion. The constant spectrum (green) was rescaled by 50% to fit into the display.
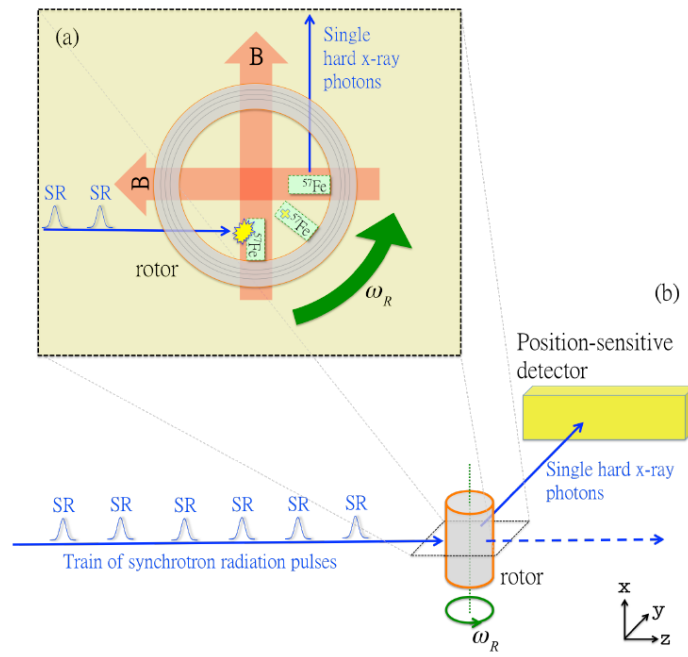
Figure 14: **a)** The Lighthouse Setup from a top view. The sample (light green rectangles) is mounted onto a small rotor. It rotates (green arrow) with frequency ω_R in and out of two static magnetic fields B (red arrows). Meanwhile, a pulse from an x-ray source, e.g. Synchrotron Radiation (SR), impinges perpendicularly on the target. **b)** The Lighthouse Setup in side view. Image courtesy of Wen-Te Liao [23].

fractions of ns [23].

SUMMARY AND OUTLOOK

In this thesis we have investigated the effects of applying periodic magnetic switching series in a Nuclear Forward Scattering (NFS) setup. We found that under the consideration of certain parameters frequency combs emerge in the forward scattered radiaton. These combs offer two possible applications. First of all, as a measurement instrument in the usual sense of frequency combs. Second, since their distinctive form can be controlled via the magnetic switching, as an information carrier.

In the NFS setup we chose (see sec. 1.2 on page 6) for our theoretical investigations, only the two $\Delta M = 0$ transitions between the ground and first excited state of ^{57}Fe Mößbauer nuclei are driven by a short x-ray pulse of 1 ns duration. When the resonant x-ray penetrates the target, the excitation is delocalized and recoilless, allowing for Multiple Scattering effects in the forward re-emitted radiation. Additionally, in the presence of an external magnetic field, the two magnetic sublevel transitions interfere and produce a pattern characteristic for NFS, the Quantum Beat (QB).
Performing a single, timed off- and subsequent on-switching of the external magnetic field permeating the nuclei leads to a coherent storage of the forward scattered photon. An inversion of the magnetic field direction however causes a π-shift of the photon phase.
These two mechanisms of storage and phase modulation constitute the basis of our periodic switching sequences. Therefore, we verified our implementation of both (chap. 3) with previously published results [1].

Our new numerical results show: Weaving a certain periodicity into the magnetic switching generates periodicity in the frequency spectrum of the forward scattered radiation. Thereby, for adequate switching parameters, equidistant frequency combs emerge in the scattered radiation.

We applied two types of switching sequences: a non-inverting and an inverting one. In the non-inverting sequence, the magnetic field B alternates between 0 and B_0, in the inverting one between $-B_0$, 0 and B_0 with $B_0 = 34.4\ T$. The two tuning parameters of our switchings are

i) the downtime τ_{off}, during which the B-field is turned off and the photon coherently stored and

ii) the uptime τ_{on}, during which the magnetic field is on and the normal QB of NFS proceeds.

To achieve maximum signal suppression, the off-switching has to be performed in a QB minimum, for which reason τ_{on} is restricted to multiples of half the QB time period T_{QB}. The storage can be chosen freely and is only limited by the natural exponential decay.

As our simulations revealed, viable frequency combs are only obtained for a minimal uptime of $\tau_{on} = T_{QB}/2$. Furthermore, the variation of the storage time alters the shape and intensity of the frequency combs: For an increasing τ_{off} the equidistant peaks narrow, manifold and drift towards the central transition frequency $\omega_0 = 14.4\ keV$ while the overall intensity declines. The peaks of the non-inverting and the inverting sequence are shifted against each other by half their equidistant separation.

Attempting a semi-analytical approach, we found that the time spectra can be modelled by a series of cosine functions combined with the exponential decay (sec. 4.3 on page 48). The corresponding analytic Fourier transform produces a series of Lorentz-curves that is in excellent agreement with the numerically solved combs.

For our results described above, we applied abrupt switching sequences with virtually infinite slopes. To give an estimate for the experimental feasibility of our research, we tested our findings with the implementation of experimentally more realistic switching sequences. For that purpose, we smoothed the switching processes. An exponential smoothing of the up and down steps allows for longer raising and declining times of the magnetic field while not losing any of the tuning flexibility. However, giving up one degree of freedom by restraining $\tau_{on} = \tau_{off}$ permits the implementation of a sinusoidal field. An experimental setup could be given by the Lighthouse Setup (see fig. 14).

Simulations with these more realistic switchings disclosed that even under an extreme smoothing in the form of the sinusoidal field, the characteristic frequency spectra stay mainly unaltered. Hence, an experimental implementation seems within the realm of feasibility.

For the advancement of this topic, we see the following possibilities.

To realize the very fast magnetic switchings, it is inevitable to further investigate experimental setups that would allow for repetitive field variations in the *ns*-range. Setups where the field is absolutely stationary, but changing relative to the sample by moving the probe are thinkable.

Another intriguing direction to explore is the use of x-ray frequency combs as information carriers. For this we would like to investigate in what fashion the combs would need to be manipulated to process information, for example in the form of logical operations as it was recently done in Ref. [20].

Part III

APPENDIX

APPENDIX A

Radiaton Gauge

In chapter 2 we used the Radiation Gauge. Let us now take a closer look at its derivation.

Given that the source of our incident radiation is virtually infinitely far from the interaction region with the sample, we may consider the fields in the free space approximation, i.e. we have no source terms in the Maxwell equations ($\mathbf{j}_r$, $\rho_r = 0$).

In this case, the Coulomb gauge is called radiation gauge and implies $\Phi = 0$.

The gauge invariance reads

$$
\begin{aligned}
\mathbf{A} &\to \mathbf{A} + \nabla \Lambda =: \mathbf{A}', \\
\Phi &\to \Phi - \partial_t \Lambda =: \Phi',
\end{aligned}
\tag{A.1}
$$

with an arbitrary function $\Lambda(\mathbf{r}, t)$.

Suppose we have $\mathbf{A}$, Φ given in the Lorentz gauge

$$
div\mathbf{A} = -\frac{1}{c^2}\partial_t\Phi.
\tag{A.2}
$$

Now, we transform as shown in Eq. (A.1) and demand $div\mathbf{A}' \overset{!}{=} 0$ (Coulomb gauge). This demand is fulfilled by setting $div\mathbf{A} = - \triangle \Lambda$.

Since the gauge cannot change the physical quantities associated with the potentials, which are the observables E and B, the Maxwell equations have to hold for both potentials: $\Box\Phi' = 4\pi\rho_r = \Box\Phi$, which implies

$$\triangle \Phi' \overset{(A.1)}{=} \triangle(\Phi - \partial_t \Lambda) = \triangle \Phi - \partial_t(\triangle \Lambda)$$

$$= \triangle \Phi + \partial_t div\mathbf{A} \overset{(A.2)}{=} \underbrace{\left(\triangle - \frac{1}{c^2}\partial_t^2\right)}_{=-\square} \Phi = -4\pi\rho_r = 0, \qquad (A.3)$$

which would make the choice of $\Phi'(\mathbf{r}, t) = \Phi'(t)$ constant over space valid. From this we may infer:

$$0 = 4\pi\rho_r = \square\Phi' = \frac{1}{c^2}\partial_t^2\Phi'(t) - \overset{=0}{\overbrace{\triangle\Phi'(t)}}$$

$$\Longleftrightarrow \partial_t^2\Phi'(t) = 0. \qquad (A.4)$$

Thus, we can choose $\Phi'(t) = \Phi'$ as constant over time as well, making it a constant in total and the choice

$$\Phi' = 0$$

valid.

This leaves for the interaction Hamiltonian:

$$\hat{H}_1 = -\frac{1}{c}\hat{\mathbf{j}}(\mathbf{r}, t)\mathbf{A}(\mathbf{r}, t). \qquad (A.5)$$

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LIST OF FIGURES

DECLARATION

I insure that I have developed and written this thesis by myself and that I have not used any means or sources other than those indicated herein.

Heidelberg, January 2017

Gregor Nicolas Ramien